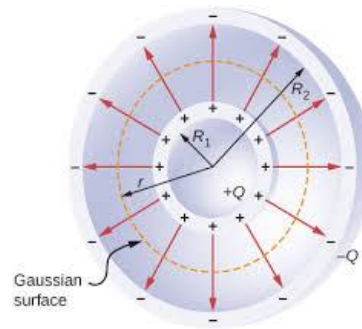


## Charged black body radiation



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$$\text{Tr} \exp(-\beta H) = \int d[\phi] \exp(iI[\phi]), \quad (3.3)$$

where the path integral is now taken over all fields which are periodic with period  $\beta$  in imaginary time. The left-hand side of (3.3) is just the parti-

Gibbons & Hawking 1977,  
Action integrals and partition  
functions in QG

What Hilbert space gives this  
background contribution ?

$$\ln Z = \underbrace{iI[g_0, \phi_0]} + \ln \int \underbrace{d[\tilde{g}] \exp(iI_2[\tilde{g}])}$$

Can one do the trace on the LHS ?

What is the simplest model ?

NB: Classical black hole  
thermodynamics comes entirely  
from background

But the normal thermodynamic argument

$$\ln Z = -WT^{-1},$$

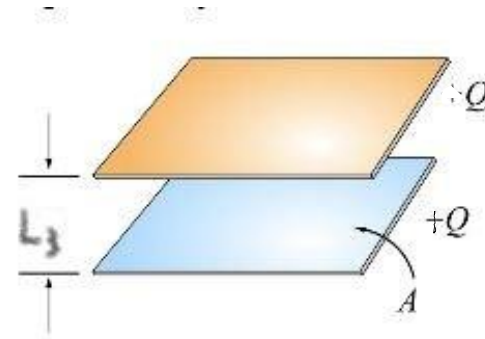
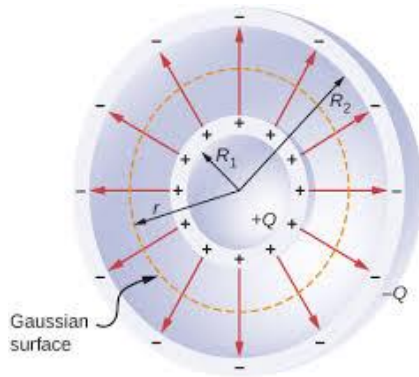
where  $W = M - TS - \sum_i \mu_i C_i$  is the “thermodynamic potential” of the system. One can therefore regard  $iI[g_0, \phi_0]$  as the contribution of the background to  $-WT^{-1}$  and the second as the contributions arising from thermal gravitons

What is the simplest model ?

Linear theory

Linearized gravity similar to  
electromagnetism

## Charged spherical or planar capacitor



$$R_1 < r < R_2$$

$$+q \quad -q$$

$$A_0 = -\phi = -\frac{q}{4\pi r}, \quad \pi^i = -\frac{qx^i}{4\pi r^3}$$

$$I = \int d^4x [\dot{A}_i \pi^i - \mathcal{H}_0 + A_0 \partial_i \pi^i]$$

$$P_1 : z = 0 \quad P_2 : z = L_3$$

$$+\frac{q}{A} \quad -\frac{q}{A}$$

$$\phi = -\frac{q}{A}z, \quad \pi^i = -\delta_3^i \frac{q}{A}$$

$$\mathcal{H}_0 = \frac{1}{2}(\pi^i \pi_i + B^i B_i)$$

Improved action with boundary term

$$I' = I + \int dt \phi_2 Q - \int dt \phi_1 Q \quad Q = - \int_S d\sigma_i \pi_L^i$$

Hawking & Ross 1995, Duality  
between electric and magnetic BHs

**Gibbons & Hawking:** contribution to partition function from Euclidean action evaluated at classical saddle point

$$\ln Z(\beta, \mu) = -I'_E + f(\beta) \quad \mu = \phi_1 - \phi_2$$

$$-I'_E = 2\pi\beta\mu^2 \frac{R_1 R_2}{R_2 - R_1}$$

$$-I'_E = \frac{\beta\mu^2 A}{2L_3}$$

Aim : microscopic explanation of semi-classical contribution

perfectly conducting boundary conditions  
as in Casimir effect

$$\vec{n} \cdot \vec{B} = 0 = \vec{n} \times \vec{E}$$

do not use heat kernel techniques but mode expansion  
possible because of very simple geometry

$$x^i = (x^a, x^3) \quad a = 1, 2 \quad k_i = \frac{\pi n_i}{L_{(i)}}$$

Dirichlet, sines only for x,y components

$$A_c(x^i) = \sum_{n_a} \sum_{n_3 > 0} A_{c,k_a,k_3}^S \sin k_3 x^3 e^{ik_a x^a}, \quad \pi^d(x^i) = \sum_{n_a} \sum_{n_3 > 0} \pi_{k_a,k_3}^{Sd} \sin k_3 x^3 e^{ik_a x^a},$$

Neumann, cosines + zero mode for z components

$$A_3(x^i) = \sum_{n_a} [\underbrace{A_{3,k_a,0}^C}_{\text{zero mode}} + \sum_{n_3 > 0} A_{3,k_a,k_3}^C \cos k_3 x^3] e^{ik_a x^a},$$

$$\pi^3(x^i) = \sum_{n_a} [\underbrace{\pi_{k_a,0}^{C3}}_{\text{zero mode}} + \sum_{n_3 > 0} \pi_{k_a,k_3}^{C3} \cos k_3 x^3] e^{ik_a x^a}.$$

2d FT!

# Planar vacuum capacitor      Edge modes

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## Detailed Hamiltonian analysis

Put the mode expansion in the canonical Hamiltonian and perform the constraint analysis **after** having taken the boundary conditions into account

$$H_c = \int d^3x \left[ \frac{1}{2} \pi^i \pi_i + \frac{1}{4} F_{ij} F^{ij} - A_0 \partial_i \pi^i \right]$$

$k_3 \neq 0$       modes related to standard black body result  
(corrected because of Casimir effect at non-zero temperature)

constraints

$$H_W = \frac{V}{2} \sum_{n_a, n_3 > 0} \left[ A_{0, k_a, k_3}^S (i k_b \pi_{b, k_a, k_3}^{*S} + k_3 \pi_{b, k_a, k_3}^{*3C}) \right]$$

standard discussion applies, only two independent polarizations

$k_3 = 0$  are physical : not affected by the constraints or by gauge transformations

$$A_i^{NPG}(x, y, 0) = \delta_i^3 \phi = \partial_i [z\phi], \quad \pi_{NPG}^i(x, y, 0) = \delta_3^i \pi = \partial_i [z\pi],$$

not longitudinal because  $z \in [0, L_3]$

$k_3 = 0$  modes give extra dynamics of massless scalar in 2+1 dimensions

$$H_{NPG} = \frac{V}{2} \sum_{n_a} \left[ \pi_{k_a,0}^{3C} \pi_{k_a,0}^{3C*} + \omega_{k_a}^2 A_{3,k_a,0}^C A_{3,k_a,0}^{C*} \right], \quad \omega_{k_a} = \sqrt{k_1^2 + k_2^2}.$$

$$S^{NPG} = \frac{L_3}{2} \int dt \int_{-L_1}^{L_1} dx \int_{-L_2}^{L_2} dy \left[ (\dot{\phi})^2 - \partial_a \phi \partial^a \phi \right].$$

Electric charge observable

$$Q = -\pi_{0,0,0}^{3C} A = - \int_{-L_1}^{L_1} dx \int_{-L_2}^{L_2} dy \pi, \quad A = 4L_1 L_2.$$



$k_a \neq 0$  modes give  $Z'_{NPG}(\beta) = \text{Tr} e^{-\beta \hat{H}'_{NPG}},$

$$\ln Z'_{NPG}(\beta) = f_A(\beta) = \frac{1}{2} b_A \beta^{-2}, \quad b_A = \frac{\zeta(3)}{\pi} A.$$

$k_a = 0$  mode corresponds to free particle

$$q = A_{3,0,0}^C \sqrt{V}, \quad p = \pi_{3,0,0}^3 \sqrt{V}$$

$$H_{NPG}^0 = \frac{1}{2} p^2, \quad Q = -\sqrt{\frac{A}{L_3}} p.$$

turn on chemical potential for electric charge  $Z_{NPG}^0(\beta, \mu) = \text{Tr} e^{-\beta \hat{H}_{NPG}^0 + \beta \mu \hat{Q}},$

$$\ln Z_{NPG}^0(\beta, \mu) = \ln \Delta q - \frac{1}{2} \ln (2\pi\beta) + \beta \mu^2 \frac{A}{2L_3},$$

reproduces the semi-classical result !

Helmholtz free energy

$$\mathcal{F}(\beta) = -\beta^{-1} \ln Z(\beta) = \mathcal{F}(0) + \mathcal{F}_2(\beta)$$

Zero temperature Casimir energy

$$\mathcal{F}(0) = \frac{1}{2} \sum_{\vec{k}} \hbar \omega_{\vec{k}} = -A \frac{\hbar c \pi^2}{720 L_3^3}$$

Thermal contribution

$$\mathcal{F}_2(\beta) = \sum_{\vec{k}} \beta^{-1} \ln (1 - e^{-\beta \hbar \omega_{\vec{k}}})$$

Subtraction of empty space BB result

$$= \frac{L^2 \pi}{d^2} \left[ -\frac{1}{2\beta} \left( \frac{d}{\pi \hbar \beta c} \right)^2 \zeta(3) + \sum_{n=1}^{\infty} b(d, T, n) + \frac{2}{\beta} \left( \frac{d}{\pi \hbar \beta c} \right)^3 \zeta(4) \right]$$

Sernelius, Surface modes

$$-\beta^{-1} f_A(\beta)$$

$$b(d, T, n) = \frac{1}{2} \frac{1}{\beta} \int_{n^2}^{\infty} ds \ln [1 - \exp(-\pi \hbar \beta c \sqrt{s}/d)]$$

# Summary

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## Main claim

**microstates** responsible for BH entropy  
related to **non-proper gauge DoF**  
rather than physical gravitons

What the **physical DoF** freedom are can  
only be decided **after** taking the  
boundary conditions (or the topology)  
into account, not before

## Further arguments

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- Gauge sector of electromagnetism as topological field theory
- Quantum Coulomb solution as coherent state of unphysical photons
- linearized Schwarzschild solution involves temporal/longitudinal DoF
- no physical gravitons in 3d but BTZ black hole
- observables = ADM surface charges involve unphysical DoF

Hamiltonian formulation

$$S = \int dt \left[ \int d^3x (\dot{A}_\mu \pi^\mu - \lambda_1 \pi^0 + A_0 \partial_i \pi^i) - H_0 \right]$$

$$H_0 = \int d^3x \left( \frac{1}{2} \pi^i \pi_i + \frac{1}{4} F^{ij} F_{ij} \right) \quad \pi^i = -E^i$$

first class constraints

$$\pi^0 = 0 = \partial_i \pi^i$$

physical DoF  $(A_i^T, \pi_T^i)$

unphysical DoF

$(A_i^L, \pi_L^i), (A_0, \pi^0)$

Quantization

reduced phase space: transverse DoF  
in positive definite Hilbert space

with charged sources: quantize transverse  
fluctuations around classical charged solution

expand transverse fields in terms of oscillators

$$[a_a(\vec{k}), a_b^\dagger(\vec{k}')] = \delta_{ab} \delta^{(3)}(\vec{k} - \vec{k}'), \quad a, b = 1, 2$$

physical Hamiltonian

$$H^{\text{phys}} = \int d^3k \, \omega_{\vec{k}} a_a^\dagger a^a, \quad \omega_{\vec{k}} = \sqrt{\vec{k} \cdot \vec{k}}$$

Partition function

$$Z(\beta) = \text{Tr} e^{-\beta H}$$
$$\ln Z(\beta) = \frac{\pi^2}{45} \beta^{-3} V$$

Drawbacks: (i) not manifestly Lorentz invariant

(ii) quantize transverse fluctuations around classical charged solution  
what is the quantum nature of this classical solution ?

quantize all polarizations in indefinite metric Hilbert space

a) Gupta-Bleuler  $[a_\mu(\vec{k}), a_\nu^\dagger(\vec{k}')] = \eta_{\mu\nu} \delta^{(3)}(\vec{k} - \vec{k}')$

physical state condition  $(\partial_\mu A^\mu)^+ |\psi\rangle^{\text{phys}} = 0$  additional (null) states decouple

b) BRST quantization  $(\mathcal{P}, C), (\bar{C}, \rho)$  spurious fermionic DoF  
cancel contributions from longitudinal and temporal photons

BRST charge  $\Omega = \int d^3x (\pi^0 \rho - \partial_i \pi^i C)$

gauge fixation  $K = \int d^3x (\bar{C} \partial_k A^k + \mathcal{P} A_0 - \frac{1}{2} \bar{C} \pi^0),$

path integral  $\int \mathcal{D}(A_\mu \pi^\mu C \mathcal{P} \rho \bar{C}) e^{iS_{\text{BRST}}^H}$

$$S_{\text{BRST}}^H = \int dt \left[ \int d^3x (\dot{A}_\mu \pi^\mu + \dot{C} \mathcal{P} + \dot{\bar{C}} \rho) - H_0 - \{\Omega, K\} \right]$$

$$\int \mathcal{D}(A_\mu C \bar{C}) e^{iS_{\text{BRST}}^L} \quad S_{\text{BRST}}^L = \int d^4x \left( \frac{1}{2} A^\nu \square A_\nu - \bar{C} \square C \right)$$

massless scalar

$$Z(\beta) = \int_{\text{periodic paths } \beta} \mathcal{D}\phi e^{-S^{\text{KG}}} = (\det \square_E)^{-\frac{1}{2}}$$

$$\ln Z(\beta) = \frac{\pi^2}{90} V \beta^{-3}$$

Hawking, CMP 55 (1977) 133

Electromagnetism: 4 massless scalars + 1 complex fermion

$$Z(\beta) = (\det \square_E)^{-2} (\det \square_E) = (\det \square_E)^{-1} \quad \longrightarrow \quad \ln Z(\beta) = \frac{\pi^2}{45} \beta^{-3} V$$



“physical” sector  $(A_i^T, \pi_T^i)$   $H^{\text{ph}} = \int d^3x \frac{1}{2} (\pi_T^i \pi_i^T - A_i^T \Delta A_T^i),$

topological sector  $(A_i^L = \partial_i A, \pi_i^L = \frac{1}{\Delta} \partial^i \pi), (A_0 \pi^0)$

$$H^{\text{gs}} = \int d^3x \mathcal{H}^{\text{gs}} = -\frac{1}{2} i \{ \Omega, \bar{\Omega} \}, \quad \bar{\Omega} = 2iK - i \int d^3x \mathcal{P} \frac{1}{\Delta} \pi$$

contains electric charge observable

expand all fields in terms of oscillators

physical transverse  $[a_a(\vec{k}), a_b^\dagger(\vec{k}')] = \delta_{ab} \delta^{(3)}(\vec{k} - \vec{k}'), \quad a, b = 1, 2$

unphysical bosonic (null)  $[a(\vec{k}), b^\dagger(\vec{k}')] = \delta^{(3)}(\vec{k} - \vec{k}')$

unphysical ghost  $[c(\vec{k}), \bar{c}^\dagger(\vec{k}')] = \delta^{(3)}(\vec{k} - \vec{k}')$

$$\eta_{\mu\nu} = \begin{pmatrix} -1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} \longrightarrow \eta_{\mu\nu} = \begin{pmatrix} 0 & 1 & 0 & 0 \\ 1 & 0 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

BRST charge  $\Omega = \int d^3k (c^\dagger a + a^\dagger c)$

physical states  $\Omega|\psi\rangle^{\text{phys}} = 0$  BRST exact states decouple  ${}^{\text{phys}}\langle\psi|\Omega|\chi\rangle = 0$

in particular  $\Omega|0\rangle = 0$  vacuum state is physical

Hamiltonian  $H_0 + [\Omega, K] = \int d^3k \omega_{\vec{k}} a_a^\dagger a^a + [\Omega, K']$

ghosts and unphysical bosonic DoF drop out

responsible for black body entropy

$$(\text{s})\text{Tr } e^{-\beta[\Omega, K']} = 0$$

Gauge sector always trivial ? No: cf. topological field theories

static charge at the origin

$$S_T = S^{EM} - \int d^4x j^\mu A_\mu, \quad j^\mu = \delta_0^\mu Q \delta^{(3)}(x)$$

only Gauss law modified

$$\partial_i \pi^i = j^0$$

modified BRST charge

$$\Omega^Q = \int d^3k [c^\dagger (a - q_{\vec{k}}) + (a^\dagger - q_{\vec{k}}) c]$$

c-number

$$q_{\vec{k}} = \frac{Q}{\sqrt{2}(2\pi)^{3/2} \omega_{\vec{k}}^{3/2}}$$

Fourier transform of

$$j^0$$

old vacuum no longer physical

$$\Omega^Q |0\rangle \neq 0$$

new vacuum

$$(a - q_{\vec{k}}) |0\rangle^Q = 0 \quad (b, c, \bar{c}, a_a) |0\rangle^Q = 0 \quad \longrightarrow \quad \Omega^Q |0\rangle^Q = 0$$

in terms of old vacuum

$$|0\rangle^Q = e^{\int d^3k q_{\vec{k}} b^\dagger(\vec{k})} |0\rangle$$

coherent state of null photons

# Electromagnetism    Quantum Coulomb solution

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$$[b(\vec{k}), b^\dagger(\vec{k}')] = 0 \quad \longrightarrow \quad \text{unusual 'classical' properties} \quad {}^Q\langle 0|0\rangle^Q = 1$$

instead of  $\langle \alpha|\beta\rangle = e^{\alpha^*\beta}$

Ehrenfest theorem

$${}^Q\langle 0|\vec{E}(x)|0\rangle^Q = -\frac{Q\vec{x}}{4\pi r^3} \qquad {}^Q\langle 0|\vec{\nabla} \times \vec{A}(x)|0\rangle^Q = 0$$

NB: requires infrared regularisation

$$FT\left(\frac{1}{k^2 + \nu^2}\right) = \frac{e^{-\nu r}}{4\pi r}, \quad \nu \rightarrow 0$$

interpretation : extrapolation of Aharonov-Bohm  
effect to quantized electromagnetic field

Dirac 1932  
Fock & Podolski 1932  
Bronstein 1936  
GB 2010

32-3\*

## On Dirac's Quantum Electrodynamics

V. FOCK AND B. PODOLSKY

Phys. Zs. Sowjetunion **1**, 798, 1932 (in English)  
Fock57, pp. 52–54

In his new paper,<sup>1</sup> Dirac suggested an original combination of the quantum electrodynamics of vacuum with the wave equation for matter. For a one-dimensional example, he demonstrated how the Coulomb interaction can appear in some approximation.

$$W\psi_2 - i\hbar \frac{\partial \psi_2}{\partial t} = \left( K - \frac{\varepsilon_1 \varepsilon_2}{4\pi |\mathbf{r}_1 - \mathbf{r}_2|} \right) \psi_0 = -U\psi_0.$$

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<sup>2</sup>We are looking for that part of the wave functional  $\varphi$  (or  $\psi$ ) that corresponds to the zero-quantum state. (This is denoted symbolically by the factor  $\delta_{j0}$ , where  $j$  is the number of light quanta.) Only the retained operator  $a_0^\dagger a_0$  sends this zero-quantum part again to a zero-quantum one, whereas the removed operators  $a_0 a_0^\dagger$ ,  $a_0^\dagger a_0^\dagger$  and  $a_0 a_0$  either cancel it or transform it into a two-quantum state. (*V. Fock*)

## QUANTUM THEORY OF WEAK GRAVITATIONAL FIELDS<sup>1</sup>

*By M. Bronstein.*

(Received on 2. January 1936)

§1. General remarks. §2. Hamiltonian form and plane waves. §3. Commutation relations and eigenvalues of the energy. §4. Let us undertake a little gedanken experiment! §5. Interaction with matter. §6. Energy transfer by gravitational waves. §7. Deduction of Newton's law of gravitation.

while in our case another commutation relation applies, namely (cf. (8))

$$[h_{00,\mathfrak{r}}, h_{00,\mathfrak{r}'}] = -\frac{h}{2\omega} \delta(\mathfrak{r} - \mathfrak{r}').$$

Neither commutation relations are introduced *ad hoc*, but originated quite naturally from the general quantum-mechanical formalism. As we shall see this suffices to obtain the correct sign of the gravitational interactions. Thus, the fundamental difference between Coulomb and Newtonian forces is explained from quantum mechanics.

Following the idea of Dirac, Fock and Podolsky<sup>7</sup> derived Coulomb's law. Our calculation proceeds exactly parallel to theirs. We start from the equations

$$\begin{aligned} \left( \frac{1}{2m_1} \mathbf{p}_1^2 + \frac{m_1}{2} h_{00}(\mathbf{r}_1) \right) \psi + \frac{h}{i} \frac{\partial \psi}{\partial t_1} &= 0, \\ \left( \frac{1}{2m_2} \mathbf{p}_2^2 + \frac{m_2}{2} h_{00}(\mathbf{r}_2) \right) \psi + \frac{h}{i} \frac{\partial \psi}{\partial t_2} &= 0. \end{aligned}$$

The sign of the right-hand side is different than in the Fock - Podolsky formula (42). When we go back to the configuration space we accordingly obtain the Schrödinger equation with the potential energy

$$-\frac{m_1 m_2}{16\pi |\mathbf{r}_1 - \mathbf{r}_2|},$$

and thus we have recovered Newtonian gravitation as a necessary consequence of the quantum theory of gravity.

The Physical-Technical Institute  
and The Physical Institute of the University.  
Leningrad, August 1935.

linearized GR = massless spin 2 gauge field on Minkowski background

**Hamiltonian formulation**  $S_{PF}[h_{mn}, \pi^{mn}, n_m, n] = \int dt \left[ \int d^3x \left( \pi^{mn} \dot{h}_{mn} - n^m \mathcal{H}_m - n \mathcal{H} \right) - H_{PF} \right],$

$$H_{PF}[h_{mn}, \pi^{mn}] = \int d^3x \left( \pi^{mn} \pi_{mn} - \frac{1}{2} \pi^2 + \frac{1}{4} \partial^r h^{mn} \partial_r h_{mn} - \frac{1}{2} \partial_m h^{mn} \partial^r h_{rn} + \frac{1}{2} \partial^m h \partial^n h_{mn} - \frac{1}{4} \partial^m h \partial_m h \right)$$

$$h_{00} = -2n \quad h_{0i} = n_i$$

$$\mathcal{H}_m = -2\partial^n \pi_{mn}, \quad \mathcal{H}_\perp = \Delta h - \partial^m \partial^n h_{mn}$$

# of comp

orthogonal  
decomposition of  
symmetric rank 2 tensor

$$\phi_{mn} = \phi_{mn}^{TT} + \phi_{mn}^T + \phi_{mn}^L,$$

$$\phi_{mn}^L = \partial_m \psi_n + \partial_n \psi_m,$$

$$\phi_{mn}^T = \frac{1}{2} (\delta_{mn} \Delta - \partial_m \partial_n) \psi^T$$

$$\phi_{mn}^{TT} = \phi_{mn} - \phi_{mn}^L - \phi_{mn}^T$$

D=4      D=3

6      3

3      2

1      1

2      0



Physical DoF: gravitational waves

ADM 1962, Dynamics of  
GR, gr-qc/0405109

canonical pairs  $(h_{mn}^{TT}(x), \pi_{TT}^{kl}(\vec{y})), (h_{mn}^L(\vec{x}), \pi_L^{kl}(y)), (h_{mn}^T(x), \pi_T^{kl}(y)).$

$$\mathcal{H}_m = 0 = \mathcal{H} \iff \pi_L^{kl} = 0 = h_{mn}^T$$

$$H^R = 0 \quad \text{D=3}$$

$$H^R = \int d^3x \left( \pi_{TT}^{mn} \pi_{mn}^{TT} + \frac{1}{4} \partial_r h_{mn}^{TT} \partial^r h_{TT}^{mn} \right). \quad \text{D=4}$$

coupling to a massive particle at rest

$$S_T = \frac{1}{16\pi} S_{PF} + \int d^4x h_{\mu\nu} T^{\mu\nu}$$

$$T^{\mu\nu} = \delta_0^\mu \delta_0^\nu M \delta^{(3)}(x)$$

only Hamiltonian constraint is affected

$$\mathcal{H}_\perp = -16\pi M \delta^{(3)}(x) \iff \begin{cases} h_{mn}^T = M \frac{x_m x_n}{r^3} \\ n = -\frac{M}{r} \end{cases} \quad \text{all other variables 0}$$

after spatial diffeo

$$h_{rr} = \frac{2M}{r} = h_{00}$$

linearized Schwarzschild solution, no TT variables involved

Quantum version in linearized gravity:  
M. Bronstein (1936)



observable      ADM mass only sees       $h_{mn}^T$

surface charge      
$$16\pi P^\perp = \oint_{S^\infty} d\sigma_m (\partial_n h_T^{mn} - \partial^m h_T) \approx M$$

related to G&H boundary term that gives non trivial value for Euclidean action

exactly like for electric charge      
$$Q = - \oint_S d\sigma_m \pi_L^m$$

# Summary

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## Main claims

- (i) Interesting edge dynamics in electromagnetism
- (ii) Related to non-proper gauge degrees of freedom

(Justified ?) hope: extendable to gravity to understand microstates for black hole entropy