



Institute for Theoretical and
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Production of massive particles from the decay of a free massless field

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Decay of a single particle state

$$\phi_0 \rightarrow \phi_1 + \phi_2 + \dots + \phi_n$$

$$\omega_i = \sqrt{p_i^2 + m_i^2}$$

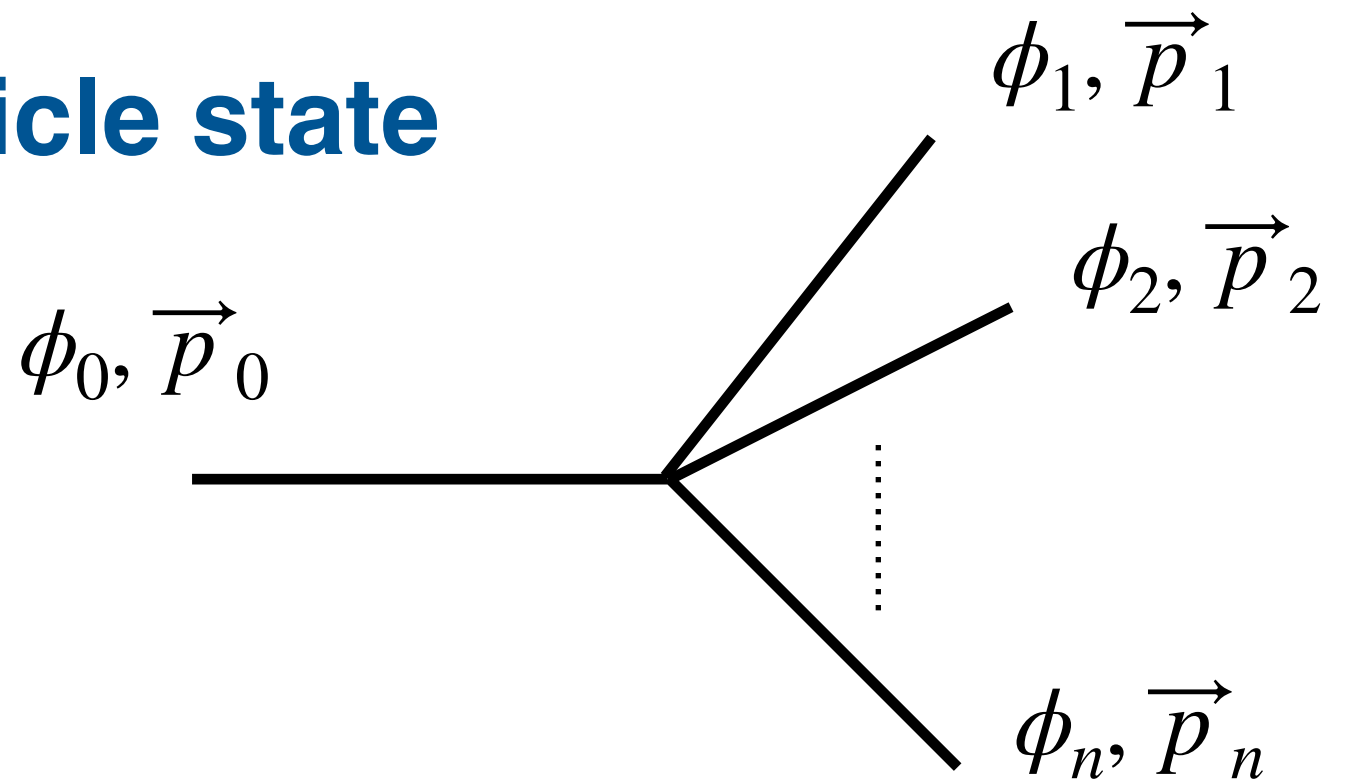
Momentum conservation:

Energy conservation:

$$\Gamma = \prod_{i=1}^n V \int \frac{d^3 k_i}{(2\pi)^3} \frac{d|S_{fi}|^2}{dt}$$

As $|i\rangle$ and $|f\rangle$ are asymptotic free states

$$S_{fi} = \frac{i(2\pi)^4}{\prod_{i=0}^n \sqrt{2\omega_i V}} \delta^4 \left(p_0 - \sum_{i=1}^n p_i \right) \mathcal{M}_{fi}$$



$$\vec{p}_0 = \vec{p}_1 + \vec{p}_2 + \dots + \vec{p}_n$$

$$\omega_0 = \omega_1 + \omega_2 + \dots + \omega_n$$

Simple toy model

$$\mathcal{H}_I = g\phi\chi^2$$

We are interested in the process $\phi \rightarrow 2\chi$

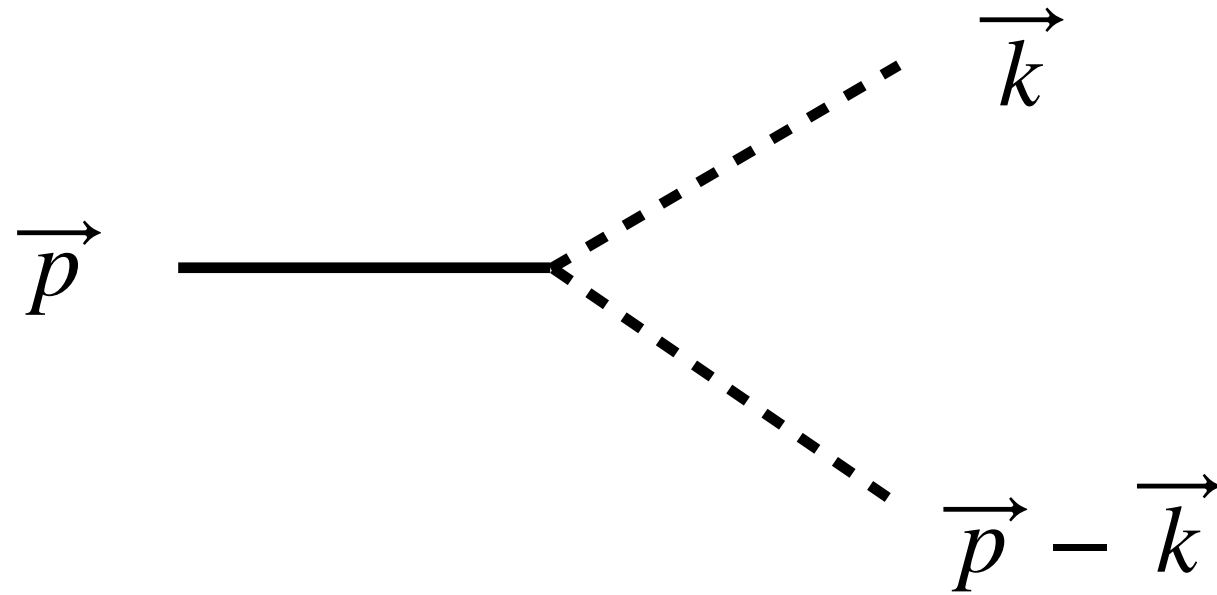
We assume that ϕ has an initial free asymptotic state with momentum $\vec{p}$

$m_\phi \longrightarrow$ **Mass of the scalar field ϕ**

$m_\chi \longrightarrow$ **Mass of the scalar field χ**

$$\omega_{\vec{k}} = \sqrt{k^2 + m_\phi^2}$$

$$\Omega_{\vec{k}} = \sqrt{k^2 + m_\chi^2}$$



we define $\epsilon_{\vec{k}} = \Omega_{\vec{k}} + \Omega_{\vec{p} - \vec{k}} - \omega_{\vec{p}}$

Two scenarios:

(I) ϕ is a one particle state

(II) ϕ is highly occupied

Two opposite cases:

(1) massive ϕ and massless χ

(2) massless ϕ and massive χ

Scenario (I)

Here final states are asymptotically free, so $\Omega_{\vec{k}}$ and $\Omega_{\vec{p}-\vec{k}}$ are the energies of the final states

$$S_{fi} = \frac{ig(2\pi)^4 \delta^4(p_\phi - k_\chi - q_\chi)}{\sqrt{2\omega_{\vec{p}}V} \sqrt{2\Omega_{\vec{k}}V} \sqrt{2\Omega_{\vec{p}-\vec{k}}V}} \quad \epsilon_{\vec{k}} = 0 \text{ is required}$$

Case (1) $\epsilon_{\vec{k}}$ can be zero

In the rest frame of the decaying particle

$$k \equiv |\vec{k}| = \frac{m_\phi}{2} \quad \Gamma_{\phi \rightarrow 2\chi} = \frac{g^2}{8\pi m_\phi}$$

Case (2) $\epsilon_{\vec{k}}$ can never be zero

$$\Gamma_{\phi \rightarrow 2\chi} = 0$$

Scenario (II)

$$\phi = \frac{\sqrt{2\langle\rho\rangle}}{\omega_{\vec{p}}} \cos\left(\vec{p} \cdot \vec{x} - \omega_{\vec{p}} t\right) \quad \langle\rho\rangle \longrightarrow \text{time averaged energy density}$$

We assume the decaying field to be highly occupied and we are not considering back reactions on it

$$\chi = \int \frac{d^3k}{(2\pi)^3} \frac{1}{\sqrt{2\Omega_{\vec{k}}}} \left(\chi_{\vec{k}}(t) e^{i\vec{k} \cdot \vec{x}} + \chi_{\vec{k}}^\dagger(t) e^{-i\vec{k} \cdot \vec{x}} \right)$$

$$[\chi_{\vec{k}}, \chi_{\vec{k}'}^\dagger] = (2\pi)^3 \delta^3(\vec{k} - \vec{k}')$$

$$A_{\vec{k}} = \chi_{\vec{k}} + \chi_{-\vec{k}}^\dagger$$

$$\alpha = \frac{g\sqrt{2\langle\rho\rangle}}{\omega_{\vec{p}}^3}$$

$$\left(\partial_t^2 + \Omega_{\vec{k}}^2\right) A_{\vec{k}} = -\omega_{\vec{p}}^2 \alpha \left(\sqrt{\frac{\Omega_{\vec{k}}}{\Omega_{\vec{k}-\vec{p}}}} A_{\vec{k}-\vec{p}} e^{-i\omega_{\vec{p}} t} + \sqrt{\frac{\Omega_{\vec{k}}}{\Omega_{\vec{k}+\vec{p}}}} A_{\vec{k}+\vec{p}} e^{i\omega_{\vec{p}} t} \right)$$

Rotating wave approximation (RWA)

$$\chi_{\vec{k}}(t) = a_{\vec{k}}(t) e^{-i\Omega_{\vec{k}} t}$$

$$\begin{aligned} H_{\chi} &= \int d^3x \left(\frac{1}{2} (\partial_t \chi)^2 + \frac{1}{2} (\vec{\nabla} \chi)^2 + \frac{1}{2} m_{\chi}^2 \chi^2 \right) \\ &= \int \frac{d^3k}{(2\pi)^3} \left(\Omega_{\vec{k}} a_{\vec{k}}^{\dagger} a_{\vec{k}} + R(t) \right) \end{aligned}$$

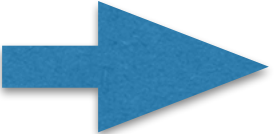
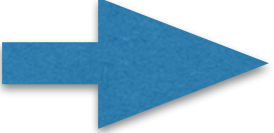
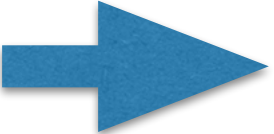
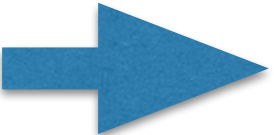
$$\begin{aligned} R(t) &= \frac{1}{2} \left(a_{\vec{k}}^{\dagger} \dot{a}_{\vec{k}} - \dot{a}_{\vec{k}}^{\dagger} a_{\vec{k}} \right) + \frac{1}{2\Omega_{\vec{k}}} \dot{a}_{\vec{k}}^{\dagger} \dot{a}_{\vec{k}} + \frac{e^{-2i\Omega_{\vec{k}} t}}{4\Omega_{\vec{k}}} \left(\dot{a}_{\vec{k}} \dot{a}_{-\vec{k}} - i\Omega_{\vec{k}} \left(\dot{a}_{\vec{k}} a_{-\vec{k}} + a_{\vec{k}} \dot{a}_{-\vec{k}} \right) \right) \\ &\quad + \frac{e^{2i\Omega_{\vec{k}} t}}{4\Omega_{\vec{k}}} \left(\dot{a}_{\vec{k}}^{\dagger} \dot{a}_{-\vec{k}}^{\dagger} + i\Omega_{\vec{k}} \left(\dot{a}_{\vec{k}}^{\dagger} a_{-\vec{k}}^{\dagger} + a_{\vec{k}}^{\dagger} \dot{a}_{-\vec{k}}^{\dagger} \right) \right) \end{aligned}$$

If $a_{\vec{k}}$ varies much slower than $e^{\pm i\Omega_{\vec{k}} t}$

$R(t)$ can be neglected, so $a_{\vec{k}}^{\dagger} a_{\vec{k}}$ is interpreted as the particle number density

$$\partial_t a_{\vec{k}} = \partial_t a_{-\vec{k}}^\dagger e^{2i\Omega_{\vec{k}} t} - i\sigma_{\vec{k}} \left(a_{\vec{k}-\vec{p}} e^{i\epsilon_{\vec{k}}^{(1)} t} + a_{\vec{p}-\vec{k}}^\dagger e^{i\epsilon_{\vec{k}}^{(2)} t} \right) \\ - i\sigma_{\vec{k}+\vec{p}} \left(a_{\vec{k}+\vec{p}} e^{i\epsilon_{\vec{k}}^{(3)} t} + a_{-\vec{p}-\vec{k}}^\dagger e^{i\epsilon_{\vec{k}}^{(4)} t} \right)$$

$$\sigma_{\vec{k}} = g \sqrt{\frac{\langle \rho \rangle / 2}{\omega_{\vec{p}}^2 \Omega_{\vec{p}-\vec{k}} \Omega_{\vec{k}}}} = \frac{\alpha \omega_{\vec{p}}^2}{2 \sqrt{\Omega_{\vec{p}-\vec{k}} \Omega_{\vec{k}}}}$$

$\epsilon_{\vec{k}}^{(1)} = \Omega_{\vec{k}} - \Omega_{\vec{k}-\vec{p}} - \omega_{\vec{p}}$		$\phi_{\vec{p}} + \chi_{\vec{k}-\vec{p}} \leftrightarrow \chi_{\vec{k}}$
$\epsilon_{\vec{k}}^{(2)} = \Omega_{\vec{k}} + \Omega_{\vec{p}-\vec{k}} - \omega_{\vec{p}}$		$\phi_{\vec{p}} \leftrightarrow \chi_{\vec{k}} + \chi_{\vec{p}-\vec{k}}$
$\epsilon_{\vec{k}}^{(3)} = \Omega_{\vec{k}} - \Omega_{\vec{k}+\vec{p}} + \omega_{\vec{p}}$		$\phi_{\vec{p}} + \chi_{\vec{k}} \leftrightarrow \chi_{\vec{k}+\vec{p}}$
$\epsilon_{\vec{k}}^{(4)} = \Omega_{\vec{k}} + \Omega_{-\vec{k}-\vec{p}} + \omega_{\vec{p}}$		vacuum $\leftrightarrow \phi_{\vec{p}} + \chi_{\vec{k}} + \chi_{-\vec{k}-\vec{p}}$

Processes with $\epsilon_{\vec{k}}^{(i)} \ll \Omega_{\vec{k}}$ are excited

$$t > \sim 0 \quad \phi_{\vec{p}} \leftrightarrow \chi_{\vec{k}} + \chi_{\vec{p}-\vec{k}}$$

$$\epsilon_{\vec{k}} \equiv \epsilon_{\vec{k}}^{(2)} \quad \partial_t a_{\vec{k}} = -i\sigma_{\vec{k}} a_{\vec{p}-\vec{k}}^{\dagger} e^{i\epsilon_{\vec{k}} t}$$

$$a_{\vec{k}}(t) = e^{i\epsilon_{\vec{k}} t} \left[a_{\vec{k}}(0) \left(\cosh(s_{\vec{k}} t) - i \frac{\epsilon_{\vec{k}}}{2s_{\vec{k}}} \sinh(s_{\vec{k}} t) \right) - i \frac{\sigma_{\vec{k}}}{s_{\vec{k}}} a_{\vec{p}-\vec{k}}(0)^{\dagger} \sinh(s_{\vec{k}} t) \right]$$

$$s_{\vec{k}} = \sqrt{\sigma_{\vec{k}}^2 - \epsilon_{\vec{k}}^2/4}$$

RWA is valid for $|s_{\vec{k}}| \ll \{\Omega_{\vec{k}}, \Omega_{\vec{p}-\vec{k}}\}$

Instability condition: $s_{\vec{k}}$ must be real

We define dimensionless parameters

$$\vec{\kappa} = \frac{\vec{k}}{\omega_{\vec{p}}} \quad \mu = \frac{m_{\chi}}{\omega_{\vec{p}}} \quad \vec{v} = \frac{\vec{p}}{\omega_{\vec{p}}}$$

$$\beta_{\vec{\kappa}} = \sqrt{\kappa^2 + \mu^2}$$

$$\eta_{\vec{k}} = \frac{s_{\vec{k}}}{\omega_{\vec{p}}} = \sqrt{\frac{\alpha^2}{4\beta_{\vec{v}-\vec{\kappa}}\beta_{\vec{\kappa}}} - \frac{1}{4}(\beta_{\vec{\kappa}} + \beta_{\vec{v}-\vec{\kappa}} - 1)^2}$$

RWA is valid for $|\eta_{\vec{\kappa}}| \ll \{\beta_{\vec{\kappa}}, \beta_{\vec{v}-\vec{\kappa}}\}$

Instability condition: $\eta_{\vec{k}}$ must be real

Case (1)

we choose $\vec{v} = 0$

Instability condition: $1/2 - \alpha < \kappa < 1/2 + \alpha$

$$\frac{m_\phi}{2} - \frac{g\sqrt{2\langle\rho\rangle}}{m_\phi^2} < k < \frac{m_\phi}{2} + \frac{g\sqrt{2\langle\rho\rangle}}{m_\phi^2}$$

$\epsilon_{\vec{k}} = 2k - m_\phi$ **it is 0 for** $k = m_\phi/2$

What about energy conservation?

We go back to the hamiltonian

$$H_\chi = \int \frac{d^3k}{(2\pi)^3} \left(\Omega_{\vec{k}} a_{\vec{k}}^\dagger a_{\vec{k}} + R(t) \right)$$

$$\begin{aligned} R(t) = & \frac{1}{2} \left(a_{\vec{k}}^\dagger \dot{a}_{\vec{k}} - \dot{a}_{\vec{k}}^\dagger a_{\vec{k}} \right) + \frac{1}{2\Omega_{\vec{k}}} \dot{a}_{\vec{k}}^\dagger \dot{a}_{\vec{k}} + \frac{e^{-2i\Omega_{\vec{k}}t}}{4\Omega_{\vec{k}}} \left(\dot{a}_{\vec{k}} \dot{a}_{-\vec{k}} - i\Omega_{\vec{k}} \left(\dot{a}_{\vec{k}} a_{-\vec{k}} + a_{\vec{k}} \dot{a}_{-\vec{k}} \right) \right) \\ & + \frac{e^{2i\Omega_{\vec{k}}t}}{4\Omega_{\vec{k}}} \left(\dot{a}_{\vec{k}}^\dagger \dot{a}_{-\vec{k}}^\dagger + i\Omega_{\vec{k}} \left(\dot{a}_{\vec{k}}^\dagger a_{-\vec{k}}^\dagger + a_{\vec{k}}^\dagger \dot{a}_{-\vec{k}}^\dagger \right) \right) \end{aligned}$$

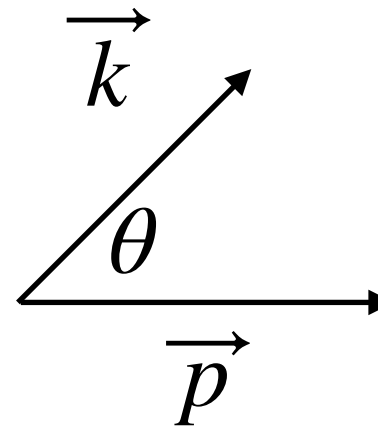
we define $\left\langle n_{\vec{k}}(t) \right\rangle = \langle 0 | a_{\vec{k}}(t) a_{\vec{k}}(t)^\dagger | 0 \rangle$

$$\langle 0 | R(t) | 0 \rangle = -\frac{1}{2} \epsilon_{\vec{k}} \left\langle n_{\vec{k}}(t) \right\rangle$$

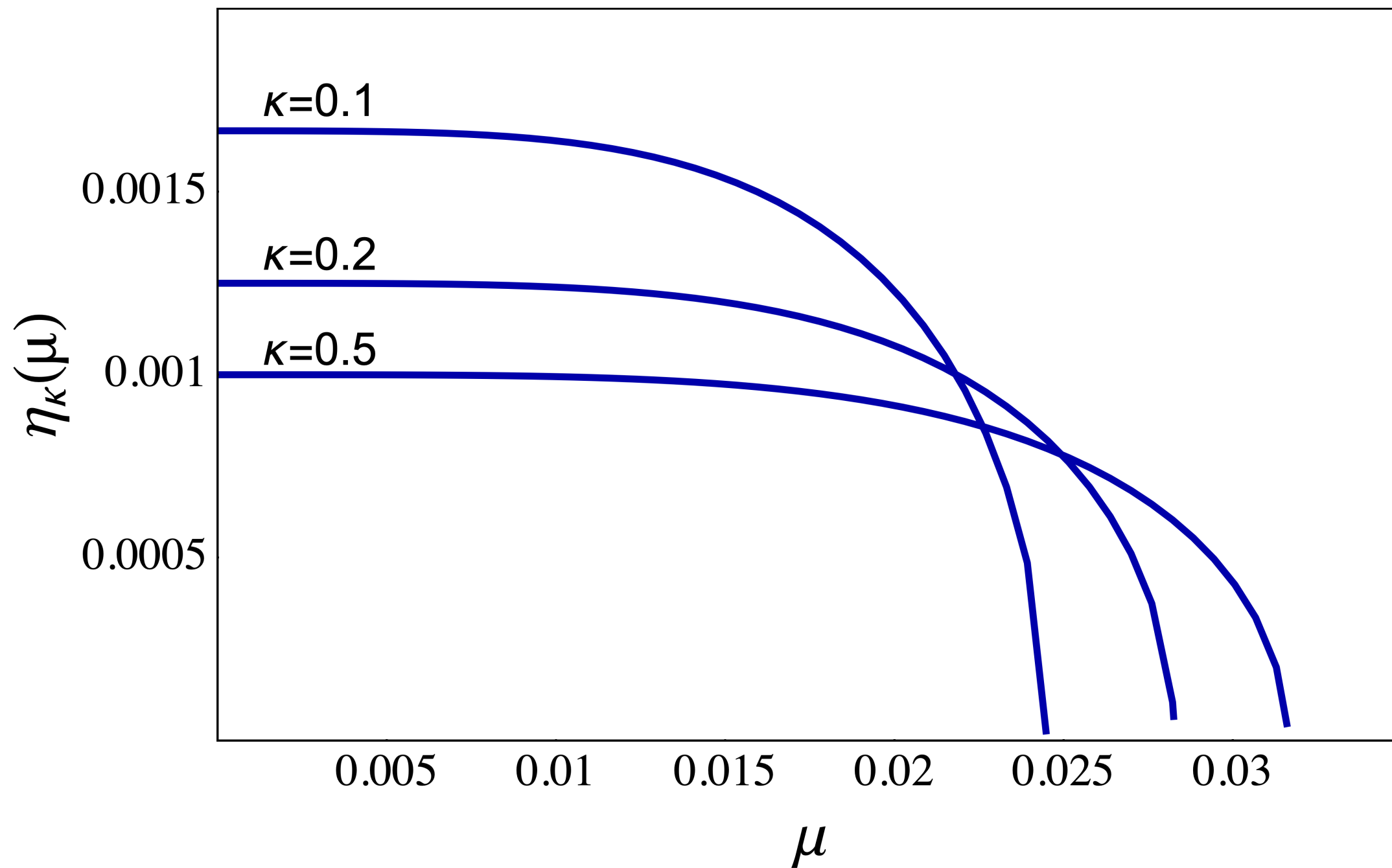
$$\left\langle E_\chi(t) \right\rangle = \frac{\omega_{\vec{p}}}{2} \int \frac{d^3k}{(2\pi)^3} \left\langle n_{\vec{k}}(t) \right\rangle$$

Case (2)

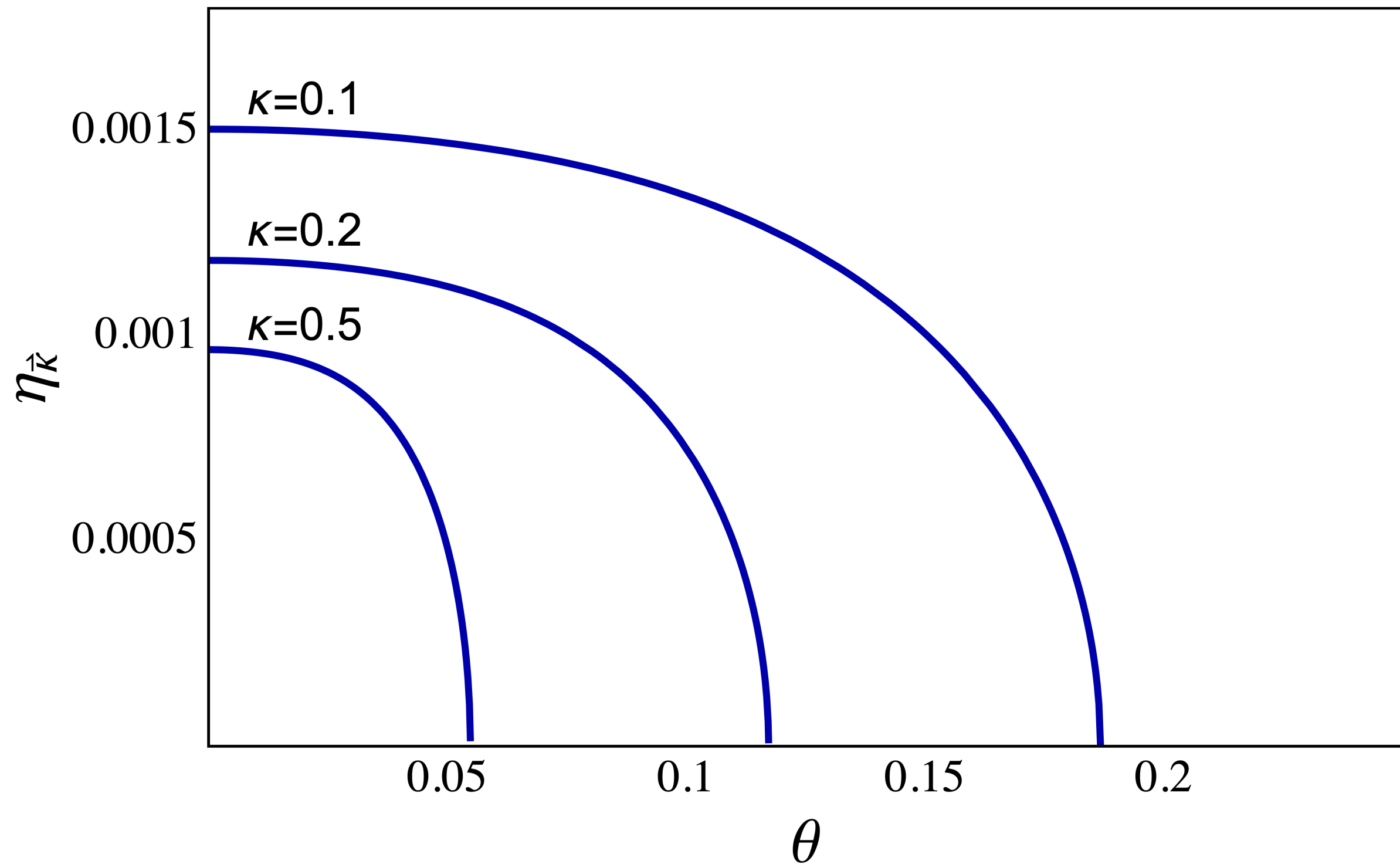
$$\nu = 1$$



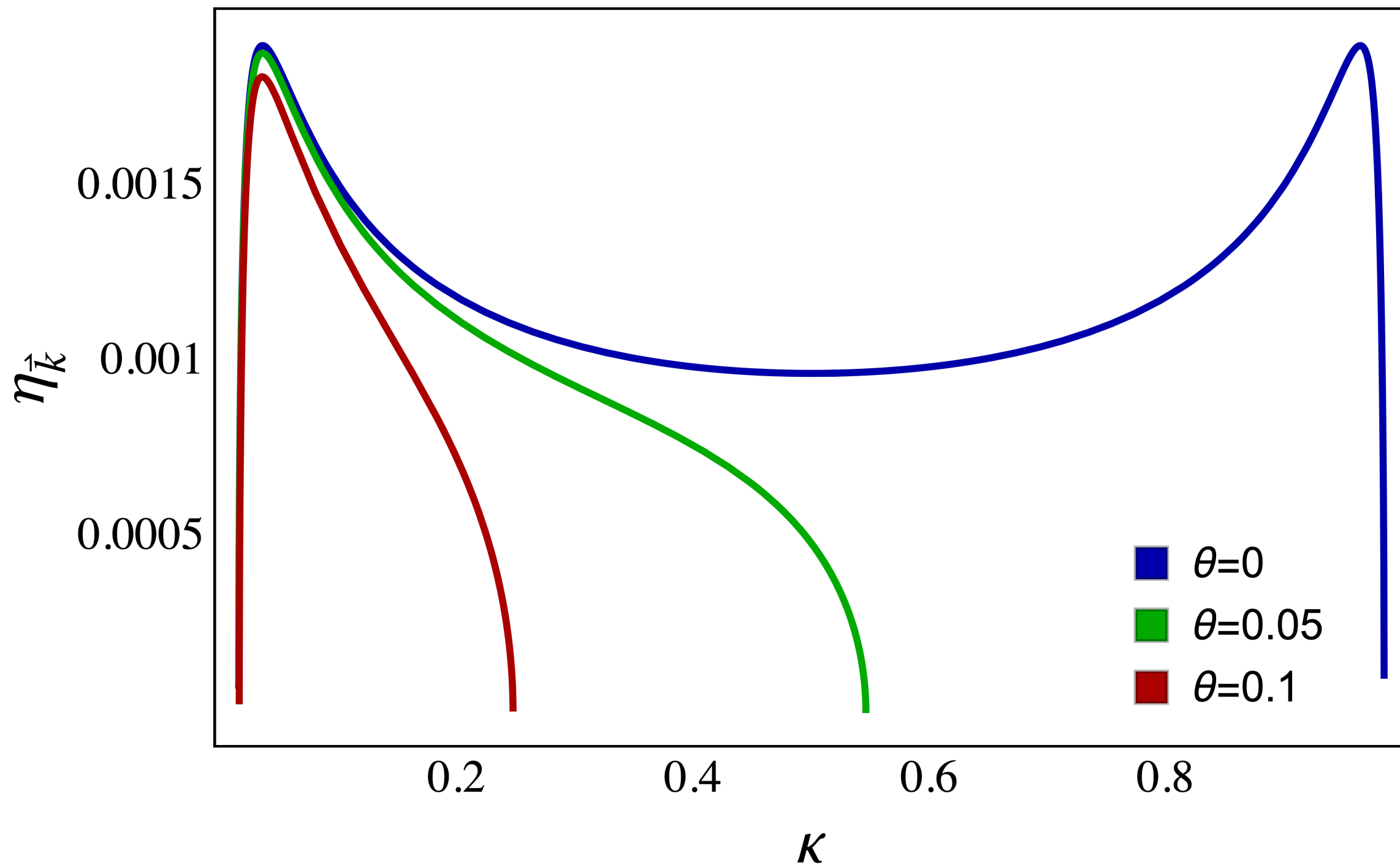
For $\theta = 0$ and $\alpha = 10^{-3}$



For $\mu = \sqrt{\alpha}/2$ **and** $\alpha = 10^{-3}$



For $\mu = \sqrt{\alpha}/2$ and $\alpha = 10^{-3}$



For $\mu \ll 1$ and $\theta \ll 1$

$$\epsilon_{\vec{k}} = \frac{p(m_\chi^2 + k^2\theta^2)}{2k(p - k)}$$

limits of the instability for $\theta = 0$

$$\kappa_- < \kappa < \kappa_+ \quad \kappa_\pm = \frac{1 \pm \sqrt{1 - \mu^4/\alpha^2}}{2}$$

$$\delta\kappa = \sqrt{1 - \mu^4/\alpha^2} \quad \Rightarrow \quad \text{Bandwidth}$$

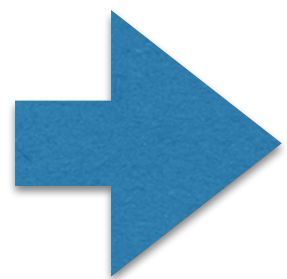
Instability condition: $\mu < \sqrt{\alpha} \quad \Rightarrow \quad \langle \rho \rangle > \frac{p^2 m_\chi^4}{2g^2}$

in case (2) the quantity $\langle \rho \rangle / p^2$ is Lorentz invariant

The RWA was used to find the instability in a frame where $p \gg m_\chi$

What about a frame where $p \ll m_\chi$?

$$\left(\partial_t^2 + \Omega_{\vec{k}}^2\right) A_{\vec{k}} = -\omega_{\vec{p}}^2 \alpha \sqrt{\frac{\Omega_{\vec{k}}}{\Omega_{\vec{k}-\vec{p}}}} A_{\vec{k}-\vec{p}} e^{-i\omega_{\vec{p}} t}$$



$$\left(\partial_t^2 + m_\chi^2\right) A_{\vec{k}} = -\omega_{\vec{p}}^2 \alpha A_{\vec{k}-\vec{p}} e^{-ip t}$$

again, the instability occurs for $\mu < \sqrt{\alpha}$

$$t \gg 0$$

when the modes $\vec{k}$ and $\vec{p} - \vec{k}$ are occupied

Case (1)

the others $\epsilon_{\vec{k}}^{(i)}$ are still big,

so there are no other processes involved

Case (2)

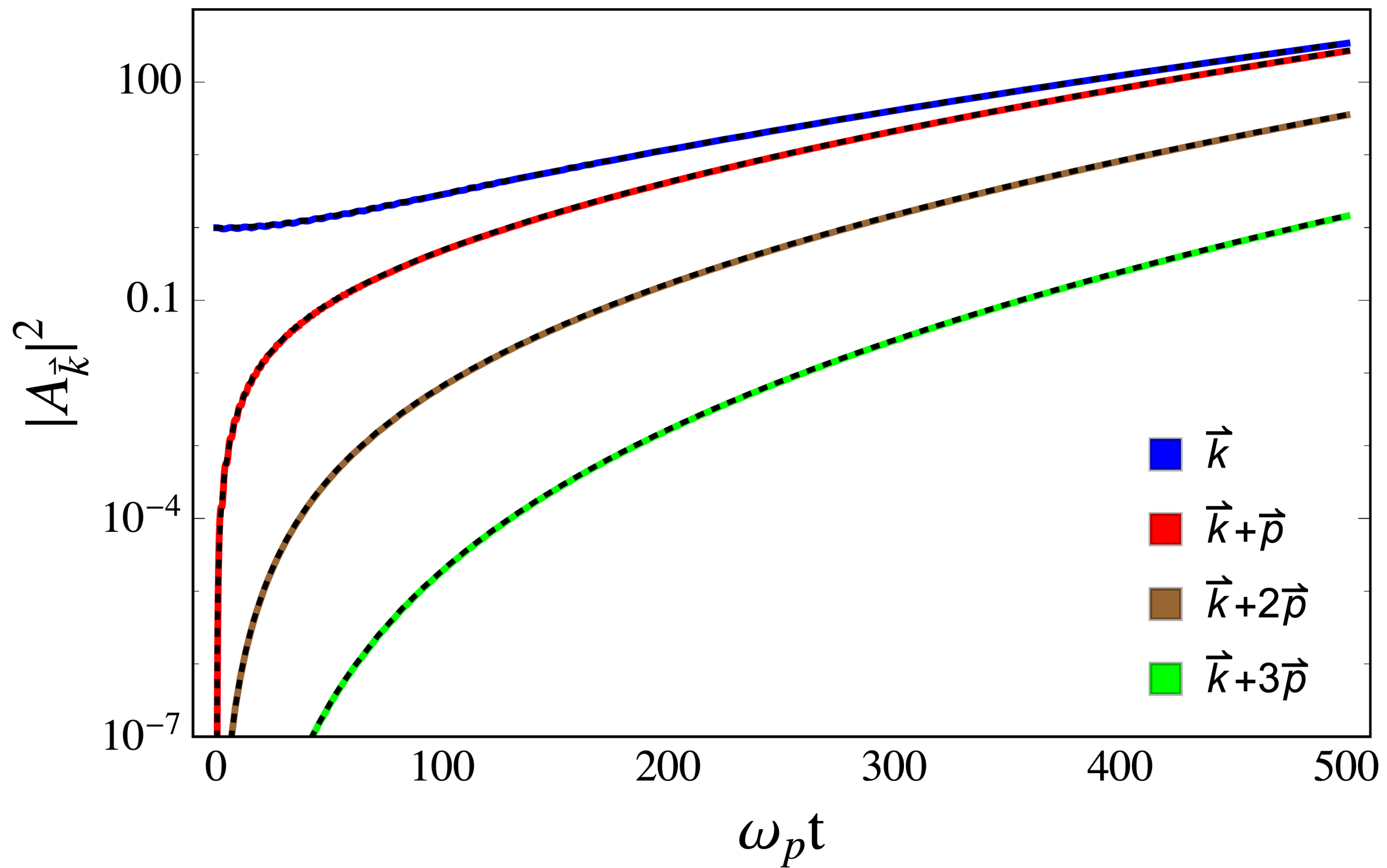
$\epsilon_{\vec{k}}^{(3)}$ is as small as $\epsilon_{\vec{k}}^{(2)}$

$$\phi_{\vec{p}} + \chi_{\vec{k} + n\vec{p}} \rightarrow \chi_{\vec{k} + (n+1)\vec{p}}$$

$$\phi_{\vec{p}} + \chi_{n\vec{p} - \vec{k}} \rightarrow \chi_{(n+1)\vec{p} - \vec{k}}$$

they gradually become important

$$\vec{k} = \vec{p}/2$$



Spontaneous decay rate

$$S_{fi} = \int_0^T dt \int d^3x \langle f | \mathcal{H}_I | i \rangle$$

$$= \frac{1}{V} \sigma_{\vec{k}} (2\pi)^3 \delta^3(\vec{k} + \vec{q} - \vec{p}) e^{i\epsilon_{\vec{k}} T/2} \frac{\sin(\epsilon_{\vec{k}} T/2)}{\epsilon_{\vec{k}}/2}$$

$$\Gamma_{\phi \rightarrow 2\chi} = V \int \frac{d^3k}{(2\pi)^3} \sigma_{\vec{k}}^2 \frac{\sin(\epsilon_{\vec{k}} T)}{\epsilon_{\vec{k}}} \quad T = \frac{1}{2s_{\vec{k}}}$$

$$\sin(\epsilon_{\vec{k}}/2s_{\vec{k}}) \approx \epsilon_{\vec{k}}/2s_{\vec{k}} \quad \Gamma_{\phi \rightarrow 2\chi} = V \int \frac{d^3k}{(2\pi)^3} \frac{\sigma_{\vec{k}}^2}{\sqrt{4\sigma_{\vec{k}}^2 - \epsilon_{\vec{k}}^2}}$$

Case (1)

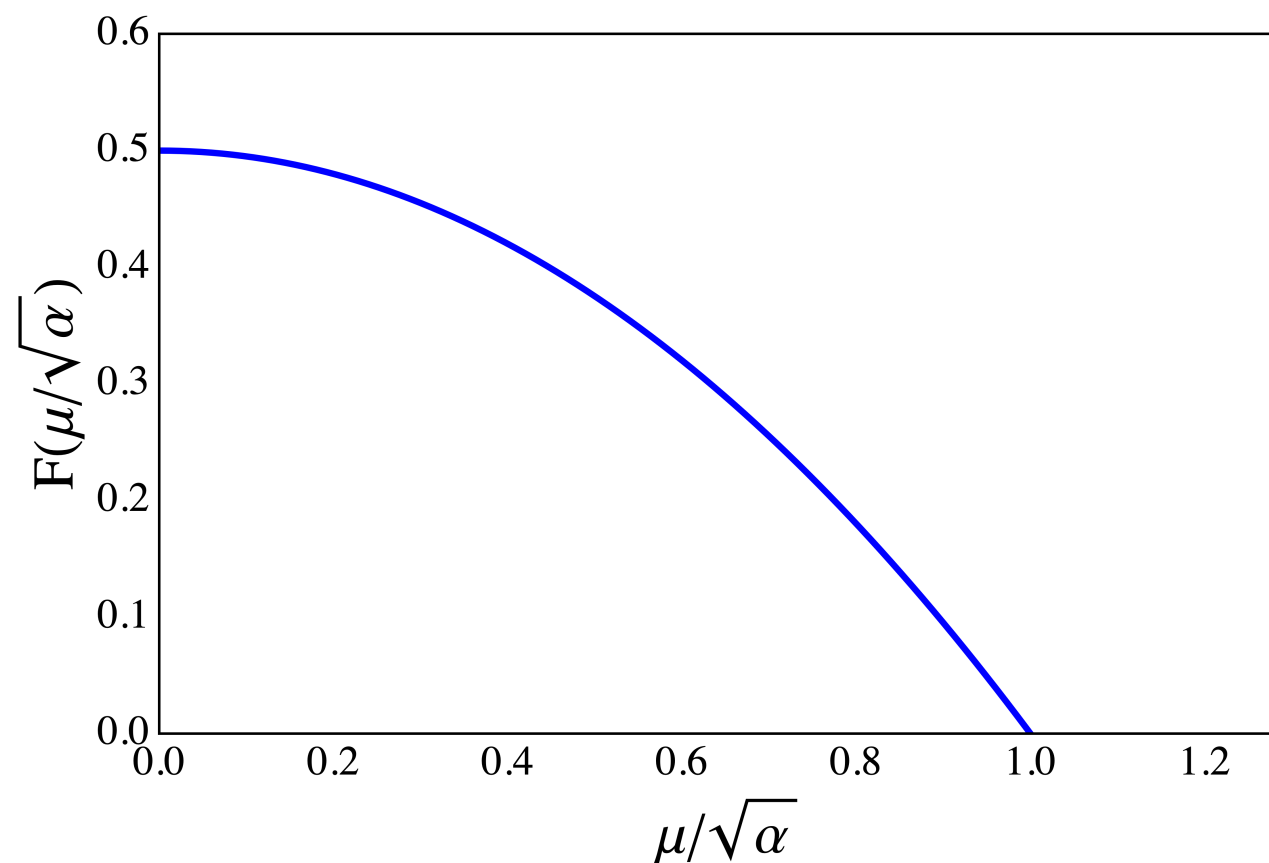
$$\Gamma_{\phi \rightarrow 2\chi}^{(1)} = \frac{g^2 N}{8\pi m_\phi}$$

Case (2)

$$\Gamma_{\phi \rightarrow 2\chi}^{(2)} = \frac{g^2 N}{8\pi p} F(\mu/\sqrt{\alpha})$$

$$F(x) = \frac{1}{\pi} \int_{\kappa_-(x)}^{\kappa_+(x)} d\kappa \left(\frac{\pi}{2} - \arcsin \left(\frac{x^2}{2\sqrt{\kappa(1-\kappa)}} \right) \right)$$

$$\frac{\Gamma_{\phi \rightarrow 2\chi}^{(2)}}{\Gamma_{\phi \rightarrow 2\chi}^{(1)}} = F(\mu/\sqrt{\alpha})$$



Thanks for your attention