

On Holographic Vacuum Misalignment

Ali Fatemiabhari

Moscow State University

Based on [arXiv:2405.08714] Daniel Elander, AF, Maurizio Piai

ITMP Seminar, MSU

May 14, 2025



Overview

- 1 Introduction
- 2 Background
- 3 The Model
 - Classes of solutions
- 4 Mass spectrum of fluctuations
- 5 Results
- 6 Summary and Outlook

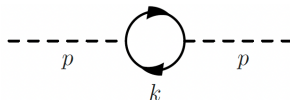
Composite Higgs Models

$$\frac{m_H^2}{\Lambda_{\text{SM}}^2} \sim 10^{-28} \lll 1 \quad \rightarrow \text{Naturalness problem}$$

- Investigations into the fundamental nature and origin of the Higgs boson are among the topical subjects in theoretical physics.
- The Higgs being a composite object with a compositeness scale of TeV order, is one of the few options for “Naturally” generating its mass.
- The Higgs boson itself, as a pseudo-Nambu-Goldstone boson associated with symmetry breaking pattern, is a reasonable candidate.

Mass terms and naturalness

$$\mathcal{L} = -\frac{1}{2}\phi(\square + m^2)\phi + \lambda\phi\bar{\psi}\psi + \bar{\psi}(i\not{\partial} - M)\psi$$



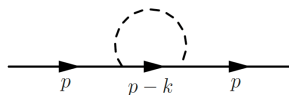
$$\Sigma_2(p^2) = \frac{3\lambda^2}{4\pi^2} \int_0^1 dx \left([M^2 - p^2 x(1-x)] \ln \frac{M^2 - p^2 x(1-x)}{\Lambda^2} + \Lambda^2 \right) + \text{finite}$$

$$m^2 = m_P^2 + \Sigma_2(p^2) \sim m_P^2 + \Lambda^2 \quad m_P = 125\text{GeV}, \Lambda \sim 10^{19}\text{GeV}$$



$$m^2 = (1 + 10^{-34}) \Lambda^2 \rightarrow \textbf{Fine-tuning (Naturalness issue)}$$

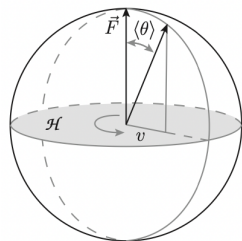
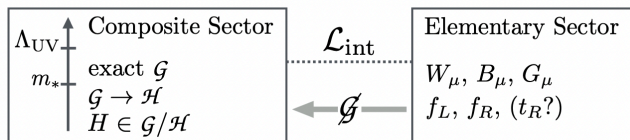
Fermion masses



$$\Sigma_2(\not{p}) = i \frac{\lambda^2}{16\pi^2} \left[\frac{\not{p} + 2M}{\varepsilon} - \int_0^1 dx [(1-x)\not{p} + M] \ln \frac{M^2 x + (1-x)(m^2 - p^2 x)}{\tilde{\mu}^2} \right]$$

$$M \rightarrow 0 \quad \Rightarrow \quad \text{Custodial chiral symmetry } \psi \rightarrow e^{i\alpha\gamma_5} \psi$$

Vacuum misalignment



$\mathcal{G} \rightarrow \mathcal{H}$ spontaneous breaking \rightarrow massless NGB's in the coset \mathcal{G}/\mathcal{H}

$$G_{EW} = SU(2)_L \times U(1)_Y \subseteq \mathcal{H}$$

G is large enough to contain at least one Higgs doublet in the coset

[Panico, Wulzer '15]

- 1 Introduction
- 2 Background
- 3 The Model**
- 4 Mass spectrum of fluctuations
- 5 Results
- 6 Summary and Outlook

Correspondance Dictionary

CHM	Gravity
D=4 ↓ Confining	D=5 ↓ D=6 compactified on a circle <i>[Witten '98]</i>
Global $SO(5)$ ↓ $SO(4)$	Local $SO(5)$ isometry ↓ Local $SO(4)$
Weakly gauged $SO(4)$ (misaligned)	Boundary terms

Instead of $SU(2) \times U(1)$, our study is gauging an $SO(4)$ group.

6D gravity action

$$\mathcal{S}_6 = \mathcal{S}_6^{(bulk)} + \sum_{i=1,2} \mathcal{S}_{5,i},$$

$$\mathcal{S}_6^{(bulk)} = \frac{1}{2\pi} \int d^6x \sqrt{-\hat{g}_6} \left\{ \frac{\mathcal{R}_6}{4} - \frac{1}{2} \hat{g}^{\hat{M}\hat{N}} (D_{\hat{M}} \mathcal{X})^T D_{\hat{N}} \mathcal{X} - \mathcal{V}_6(\mathcal{X}) \right. \\ \left. - \frac{1}{2} \text{Tr} \left[\hat{g}^{\hat{M}\hat{P}} \hat{g}^{\hat{N}\hat{Q}} \mathcal{F}_{\hat{M}\hat{N}} \mathcal{F}_{\hat{P}\hat{Q}} \right] \right\},$$

$\rho_1 < \rho < \rho_2$ and the space-time index is $\hat{M} = 0, 1, 2, 3, 5, 6$.

$$\mathcal{V}_6(\phi) = -5 - \frac{\Delta(5-\Delta)}{2} \phi^2 - \frac{5\Delta^2}{16} \phi^4,$$

$$\mathcal{X} = \exp \left[2i \sum_{\hat{A}} \pi^{\hat{A}} t^{\hat{A}} \right] \mathcal{X}_0 \phi, \quad \text{where} \quad \mathcal{X}_0 = (0, 0, 0, 0, 1)^T,$$

with $\hat{A} = 1, \dots, 4$, labelling the generators of the $SO(5)/SO(4)$ coset. For the gauge field and the scalar we have:

$$10 \rightarrow 6 \oplus 4 \rightarrow 3 \oplus 3 \oplus 3 \oplus 1, \quad 5 \rightarrow 4 \oplus 1 \rightarrow 3 \oplus 1 \oplus 1.$$

Asymptotically in the UV, the dual field theory flows towards a CFT in 5 dimensions, deformed by the insertion of an operator \mathcal{O} with scaling dimension given by $\max(\Delta, 5 - \Delta)$.

The two parameters appearing in the solution of the corresponding second-order classical equations correspond in field-theory terms to the coupling and condensate associated with \mathcal{O} .

Asymptotics and dimensional reduction to five dimensions

$$ds_6^2 = e^{-2\chi} dx_5^2 + e^{6\chi} (d\eta + \chi_M dx^M)^2,$$

where the space-time index is $M = 0, 1, 2, 3, 5$.

A part of the reduced action then becomes

$$\begin{aligned} \mathcal{S}_5^{1(bulk)} = \int d^5x \sqrt{-g_5} \left\{ \frac{R}{4} - \frac{1}{2} g^{MN} \left[6 \partial_M \chi \partial_N \chi + \partial_M \phi \partial_N \phi \right] - e^{-2\chi} \mathcal{V}_6(|\phi|) \right. \\ \left. - \frac{1}{16} e^{8\chi} g^{MP} g^{NQ} F_{MN}^{(\chi)} F_{PQ}^{(\chi)} \right\}, \end{aligned}$$

We consider background solutions in which $\chi_M = 0$, while the metric g_{MN} , ϕ , and χ depend on the radial coordinate only. The metric in five dimensions takes the domain-wall (DW) form

$$ds_5^2 = dr^2 + e^{2A(r)} dx_{1,3}^2 = e^{2\chi(\rho)} d\rho^2 + e^{2A(\rho)} dx_{1,3}^2,$$

with $d\rho = e^{-\chi} dr$.

UV expansions

$$\rho \rightarrow +\infty, \quad \phi = 0 \quad \text{critical point of } \mathcal{V}_6, \quad \chi \simeq \frac{1}{3}\rho, \quad A \simeq \frac{4}{3}\rho$$

We classify all the solutions in terms of a power expansion in the small parameter $z \equiv e^{-\rho}$. The expansion depends on five free parameters.

$$\begin{aligned} \phi(z) &= \phi_J z^{\Delta_J} + \cdots + \phi_V z^{\Delta_V} + \cdots, \\ \chi(z) &= \chi_U - \frac{1}{3} \log(z) + \cdots + (\chi_5 + \cdots) z^5 + \cdots, \\ A(z) &= A_U - \frac{4}{3} \log(z) + \cdots. \end{aligned}$$

The two parameters appearing in the solution of the corresponding second-order classical equations correspond in field-theory terms to the coupling and condensate associated with \mathcal{O} .

For $\Delta = 3$:

$$\phi(z) = \phi_J z^2 + \phi_V z^3 - \frac{25}{48} \phi_J^3 z^6 - \frac{57}{80} \phi_J^2 \phi_V z^7 + \mathcal{O}(z^8) ,$$

$$\chi(z) = \chi_U - \frac{1}{3} \log(z) - \frac{1}{24} \phi_J^2 z^4 + \left(\chi_5 - \frac{2}{25} \phi_J \phi_V \right) z^5 - \frac{1}{24} \phi_V^2 z^6 + \mathcal{O}(z^8) ,$$

$$A(z) = A_U - \frac{4}{3} \log(z) - \frac{1}{6} \phi_J^2 z^4 + \left(\frac{1}{4} \chi_5 - \frac{8}{25} \phi_J \phi_V \right) z^5 - \frac{1}{6} \phi_V^2 z^6 + \mathcal{O}(z^8) .$$

Background solutions

Confining solutions The confining solutions are such that the circle parametrised by η shrinks to zero size at some point ρ_o of the radial direction ρ and there is no conical singularity. For small $(\rho - \rho_o)$, we find that such solutions have the following form

$$\begin{aligned}\phi(\rho) &= \phi_I - \frac{1}{16} \Delta \phi_I (20 + \Delta (5\phi_I^2 - 4)) (\rho - \rho_o)^2 + \mathcal{O}((\rho - \rho_o)^2) , \\ \chi(\rho) &= \chi_I + \frac{1}{3} \log(\rho - \rho_o) + \mathcal{O}((\rho - \rho_o)^4) , \\ A(\rho) &= A_I + \frac{1}{3} \log(\rho - \rho_o) + \mathcal{O}((\rho - \rho_o)^2) ,\end{aligned}$$

Singular domain-wall solutions They obey the DW ansatz $A = 4\chi = \frac{4}{3}\mathcal{A}$. In six dimensions, they take the form of Poincaré domain walls.

$$\begin{aligned}\phi(\rho) &= \phi_I - \sqrt{\frac{2}{5}} \log(\rho - \rho_o) + \mathcal{O}((\rho - \rho_o)^2) , \\ \mathcal{A}(\rho) &= \frac{1}{5} \log(\rho - \rho_o) + \mathcal{O}((\rho - \rho_o)^2) .\end{aligned}$$

The system of equations is symmetric under the change $\phi \rightarrow -\phi$, hence a second branch of solutions can be obtained by just changing the sign of ϕ . **These solutions are singular at the end of space.**

Boundary-localised interactions

We add to the bulk action, $S_5^{(bulk)}$, several boundary terms—denoted as in order to obtain the desired five-dimensional action,

$$S_5 = S_5^{(bulk)} + \sum_{i=1,2} \left(\mathcal{S}_{\text{GHY},i} + \mathcal{S}_{\lambda,i} \right) + \mathcal{S}_{P_5,2} + \mathcal{S}_{\mathcal{V}_4,2} + \mathcal{S}_{\mathcal{A},2} + \mathcal{S}_{\mathcal{X},2} + \mathcal{S}_{\mathcal{X},2},$$

$$\mathcal{S}_{P_5,2} = \int d^4x \sqrt{-\tilde{g}} \left\{ -\frac{1}{2} K_5 \tilde{g}^{\mu\nu} (D_\mu P_5) D_\nu P_5 - \lambda_5 (P_5^T P_5 - v_5^2)^2 \right\} \Big|_{\rho=\rho_2},$$

$$\mathcal{S}_{\mathcal{V}_4,2} = - \int d^4x \sqrt{-\tilde{g}} \mathcal{V}_4(\mathcal{X}, \chi, P_5) \Big|_{\rho=\rho_2},$$

$$\mathcal{S}_{\mathcal{A},2}|_{P_5=\bar{P}_5} = \int d^4x \sqrt{-\tilde{g}} \left\{ -\frac{1}{4} \hat{D}_2 \tilde{g}^{\mu\rho} \tilde{g}^{\nu\sigma} \mathcal{F}_{\mu\nu}^{\hat{A}} \mathcal{F}_{\rho\sigma}^{\hat{A}} - \frac{1}{4} \bar{D}_2 \tilde{g}^{\mu\rho} \tilde{g}^{\nu\sigma} \mathcal{F}_{\mu\nu}^{\bar{A}} \mathcal{F}_{\rho\sigma}^{\bar{A}} \right\} \Big|_{\rho=\rho_2}.$$

$$\mathcal{S}_{\mathcal{X},2} = \int d^4x \sqrt{-\tilde{g}} \left\{ -\frac{1}{2} K_{\mathcal{X},2} \tilde{g}^{\mu\nu} (D_\mu \mathcal{X})^T D_\nu \mathcal{X} \right\} \Big|_{\rho=\rho_2}.$$

Boundary term	Parameters	Free parameters
$\mathcal{S}_{\text{GHY},i}$	—	—
$\mathcal{S}_{\lambda,i}$	$m_{\mathcal{X},i}^2$	—
$\mathcal{S}_{\mathcal{X},2}$	$D_{\mathcal{X},2}$	—
$\mathcal{S}_{P_5,2}$	$K_5(k_5), \lambda_5, v_5$	—
$\mathcal{S}_{\mathcal{V}_4,2}$	$\partial_v^2 \mathcal{V}_4(v, m_4^2)$	v, m_4^2
$\mathcal{S}_{\mathcal{A},2}$	$\bar{D}_2(\bar{\varepsilon}), \hat{D}_2(\hat{\varepsilon})$	$\bar{\varepsilon}$
$\mathcal{S}_{\mathcal{X},2}$	$K_{\mathcal{X},2}(k_{\mathcal{X}})$	$k_{\mathcal{X}}$

- 1 Introduction
- 2 Background
- 3 The Model
- 4 Mass spectrum of fluctuations**
- 5 Results
- 6 Summary and Outlook

Linearised equations for scalar fluctuations

We write here the action of the two-derivative sigma-model in D dimensions, consisting of n scalars Φ^a , with $a = 1, \dots, n$, coupled to gravity:

$$\mathcal{S} = \int d^D x \sqrt{-g} \left[\frac{R}{4} - \frac{1}{2} g^{MN} G_{ab} \partial_M \Phi^a \partial_N \Phi^b - \mathcal{V}(\Phi^a) \right]. \quad (1)$$

Where $M = 0, \dots, 3, 5, \dots, D$ the D -dimensional space-time indexes.

If we adopt the ansatz for the background solutions, that the metric is of the DW form, and that the scalar fields only depend on the radial coordinate,

$$ds_D^2 = dr^2 + e^{2A(r)} dx_{1,D-2}^2, \quad (2)$$

$$\Phi^a = \Phi^a(r), \quad (3)$$

The fluctuations around the classical background of the active scalars and gravity are treated with the gauge-invariant formalism [Bianchi et al. '20, Berg et al. '20, Elander '20, Elander, Piai '20 ...]:

$$\Phi^a(x^\mu, r) = \Phi^a(r) + \varphi^a(x^\mu, r),$$

By decomposing the metric according to the ADM formalism, one writes

$$ds_D^2 = ((1 + \nu)^2 + \nu_\sigma \nu^\sigma) dr^2 + 2\nu_\mu dx^\mu dr + e^{2A(r)} (\eta_{\mu\nu} + h_{\mu\nu}) dx^\mu dx^\nu,$$

and

$$h^\mu{}_\nu = \epsilon^\mu{}_\nu + iq^\mu \epsilon_\nu + iq_\nu \epsilon^\mu + \frac{q^\mu q_\nu}{q^2} H + \frac{1}{D-2} \delta^\mu{}_\nu h,$$

where $\epsilon^\mu{}_\nu$ is transverse and traceless, ϵ^μ is transverse.

We form the following gauge-invariant (under diffeomorphisms) combinations:

$$\mathfrak{a}^a = \varphi^a - \frac{\partial_r \Phi^a}{2(D-2)\partial_r A} h,$$

$$\mathfrak{b} = \nu - \partial_r \left(\frac{h}{2(D-2)\partial_r A} \right),$$

$$\mathfrak{c} = e^{-2A} \partial_\mu \nu^\mu - \frac{e^{-2A} q^2 h}{2(D-2)\partial_r A} - \frac{1}{2} \partial_r H,$$

$$\mathfrak{d}^\mu = e^{-2A} P^\mu{}_\nu \nu^\nu - \partial_r \epsilon^\mu,$$

For the scalar fluctuations $\alpha^a = \alpha^a(q, \rho)$, where q^μ is the four-momentum:

$$\left[\partial_\rho^2 + (4\partial_\rho A - \partial_\rho \chi) \partial_\rho - e^{2\chi-2A} q^2 \right] \alpha^a - e^{2\chi} \mathcal{X}_c^a \alpha^c = 0.$$

With $G_{ab} = \text{diag}(1, 6)$, \mathcal{X}_c^a reads as follows:

$$\begin{aligned} \mathcal{X}_c^a \equiv & \frac{\partial}{\partial \Phi^c} \left(G^{ab} \frac{\partial(e^{-2\chi} \mathcal{V}_6)}{\partial \Phi^b} \right) + \frac{4}{3\partial_\rho A} \left[\partial_\rho \Phi^a \frac{\partial(e^{-2\chi} \mathcal{V}_6)}{\partial \Phi^c} + G^{ab} \frac{\partial(e^{-2\chi} \mathcal{V}_6)}{\partial \Phi^b} \partial_\rho \Phi^d G_{dc} \right] \\ & + \frac{16(e^{-2\chi} \mathcal{V}_6)}{9(\partial_\rho A)^2} \partial_\rho \Phi^a \partial_\rho \Phi^b G_{bc}. \end{aligned}$$

A discrete spectrum can be found by imposing the following boundary conditions:

$$e^{-2\chi} \partial_\rho \Phi^c \partial_\rho \Phi^d G_{db} \partial_\rho \alpha^b \Big|_{\rho_i} = \left[\frac{3\partial_\rho A}{2} e^{-2A} q^2 \delta_b^c + \partial_\rho \Phi^c \left(\frac{4\mathcal{V}_6}{3\partial_\rho A} \partial_\rho \Phi^d G_{db} + \frac{\partial \mathcal{V}_6}{\partial \Phi^b} \right) \right] \alpha^b \Big|_{\rho_i},$$

Physical composite states in the dual theory have mass $M^2 = -q^2$.

The equations of motion for the (transverse and traceless) tensor fluctuations ϵ^μ_ν :

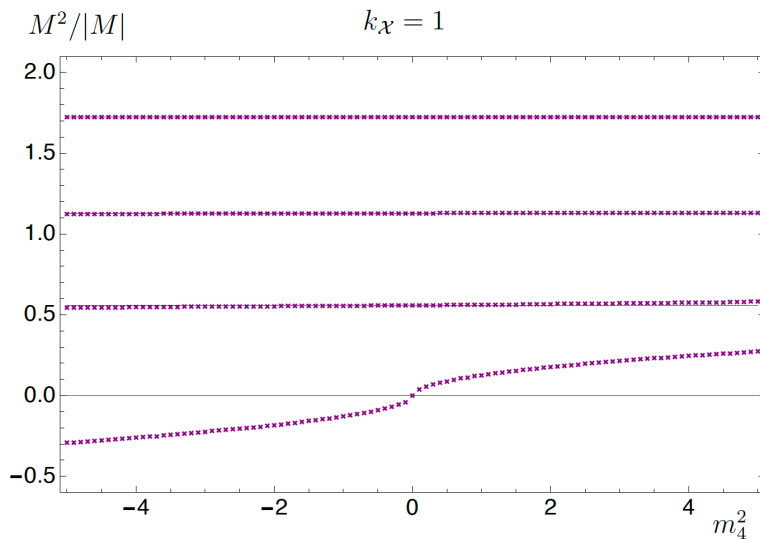
$$0 = \left[\partial_\rho^2 + (4\partial_\rho A - \partial_\rho \chi) \partial_\rho - e^{2\chi-2A} q^2 \right] \epsilon^\mu_\nu,$$

For the vector χ_M one looks at the gauge-invariant transverse polarisations,

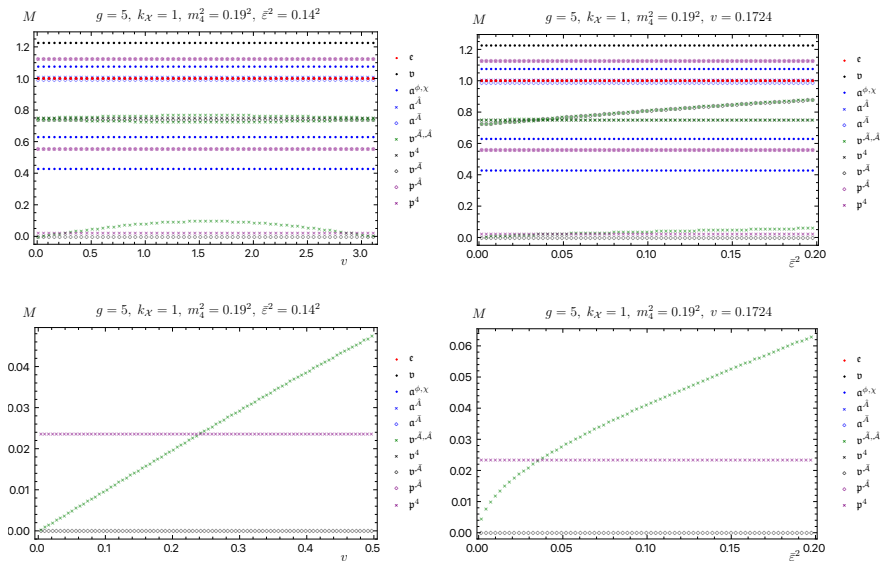
$$0 = P^{\mu\nu} \left[\partial_\rho^2 + (2\partial_\rho A + 7\partial_\rho \chi) \partial_\rho - e^{2\chi-2A} q^2 \right] \chi_\nu,$$

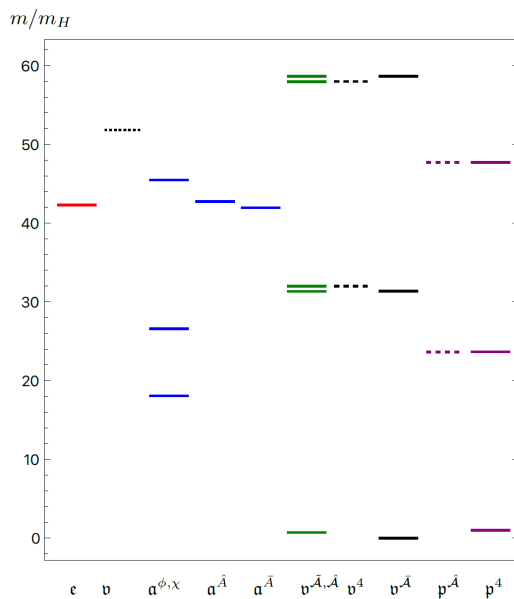
where $P^{\mu\nu} \equiv \eta^{\mu\nu} - \frac{q^\mu q^\nu}{q^2}$. Neumann boundary conditions is imposed on these two kinds of fluctuations.

Results



Results





Summary and Outlook

- We have studied a bottom-up model, with a completely smooth gravity background, that implements a simple realisation of the holographic description of confinement in the dual gauge theory.
- We presented the mass spectrum of bosonic states in the field theory.
- All the new particles are parametrically heavy with respect to the bosons that play the role of the Z , W , and Higgs boson
- We can further extend this six-dimensional model to more realistic composite Higgs models with gauged $SU(2) \times U(1)$ and fermions.
- Extension to a top-down model is expected.

Thank you