

Stress tensor and conformal correlators

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Introduction

In CFTs in d spacetime dimensions an important set of operators are *primary* operators which are annihilated by special conformal transformations, $[K_\mu, \mathcal{O}_k] = 0$.

They have definite *scaling* under spacetime dilatations $x^\mu \rightarrow \lambda x^\mu$ generated by the dilatation operator D , $[D, \mathcal{O}_k] = \Delta_k \mathcal{O}_k$.

All other operators are *descendants* of the primaries, of the type $\partial^n \mathcal{O}_k$.

Introduction

Correlators (2- and 3- point functions) of primaries are fixed by Ward identities:

$$\langle \mathcal{O}_i(x) \mathcal{O}_j(0) \rangle = \frac{\delta_{ij}}{|x|^{2\Delta_i}}$$

$$\langle \mathcal{O}_i(x_1) \mathcal{O}_j(x_2) \mathcal{O}_k(x_3) \rangle = \frac{\lambda_{\mathcal{O}_i \mathcal{O}_j \mathcal{O}_k}}{|x_{12}|^{\Delta_{ijk}} |x_{13}|^{\Delta_{ikj}} |x_{23}|^{\Delta_{jki}}}$$

where $\Delta_{ijk} = \Delta_i + \Delta_j - \Delta_k$ and $\lambda_{\mathcal{O}_i \mathcal{O}_j \mathcal{O}_k}$ are the OPE coefficients,

$$\mathcal{O}_i(x) \mathcal{O}_j(0) = \sum_k \lambda_{\mathcal{O}_i \mathcal{O}_j \mathcal{O}_k} \frac{\mathcal{O}_k(0)}{|x|^{\Delta_{ijk}}}$$

Introduction

Four point functions depend on z, \bar{z} (the positions of operator insertions) and admit a conformal block $[\mathcal{B}_k(z, \bar{z})]$ decomposition:

$$\langle \mathcal{O}_1(\infty) \mathcal{O}_2(1) \mathcal{O}_2(z, \bar{z}) \mathcal{O}_1(0) \rangle = \sum_k \lambda_{\mathcal{O}_1 \mathcal{O}_1 \mathcal{O}_k} \lambda_{\mathcal{O}_2 \mathcal{O}_2 \mathcal{O}_k} \mathcal{B}_k(z, \bar{z})$$

where $\mathcal{B}_k(z, \bar{z})$ are known functions (Dolan, Osborn) which contain contributions of the primary operator \mathcal{O}_k and its descendants.

Introduction

Central charge C_T appears in the 2-point function of stress tensors,

$$\langle T_{\mu\nu}(x) T_{\alpha\beta}(0) \rangle = \frac{C_T}{x^{2d}} \left(\frac{1}{2} I_{\mu\alpha} I_{\nu\beta} + \frac{1}{2} I_{\mu\beta} I_{\nu\alpha} - \frac{1}{d} \delta_{\mu\nu} \delta_{\alpha\beta} \right)$$

where $I_{\mu\nu} = \delta_{\mu\nu} - 2 \frac{x_\mu x_\nu}{x^2}$. We will consider large- N CFTs with $N^2 \sim C_T \gg 1$.

Major players are double-trace operators, $[OO]_{n,\ell} \simeq \mathcal{O} \partial^\ell \square^n \mathcal{O}$. For unit-normalized operators,

$$\lambda_{\mathcal{O}_1 \mathcal{O}_1 [\mathcal{O}_1 \mathcal{O}_1]_{n,\ell}} \sim 1, \quad \lambda_{\mathcal{O}_1 \mathcal{O}_1 [\mathcal{O}_2 \mathcal{O}_2]_{n,\ell}} \sim \frac{1}{C_T}$$

Introduction

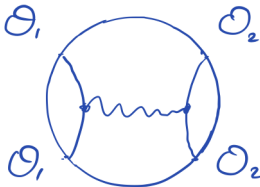
Holography relates certain large -N CFTs in d dimensions to theories of gravity in $d + 1$ dimensions.

$$\langle e^{\int d^d x J(x) \mathcal{O}(x)} \rangle = Z_{gravity}[\phi(z, x) \rightarrow J(x)]$$

For $\mathcal{O} = T_{\mu\nu}$, the source is the metric "on the boundary of AdS",
 $J(x) = g_{\mu\nu}$.

Introduction

In holographic CFTs the leading $1/C_T \sim 1/N^2$ connected four-point function receives a contribution from a Witten diagram with the graviton exchange.



Introduction

This diagram admits a conformal block decomposition, starting with the (unit-normalized) stress tensor, $T_{\mu\nu}$. The OPE coefficient $\lambda_{\mathcal{O}\mathcal{O}T_{\mu\nu}} \sim \frac{\Delta}{\sqrt{C_T}}$ is fixed by a Ward identity, so the stress-tensor contributes $\Delta_1\Delta_2/C_T \sim 1/N^2$.

In addition, there are conformal blocks of double trace operators $[\mathcal{O}_1\mathcal{O}_1]_{n,l}$ and $[\mathcal{O}_2\mathcal{O}_2]_{n,l}$, with known OPE coefficients, which also contribute

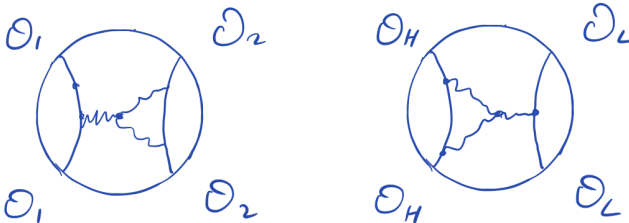
$$\lambda_{\mathcal{O}_1\mathcal{O}_1[\mathcal{O}_1\mathcal{O}_1]_{n,l}}\lambda_{\mathcal{O}_1\mathcal{O}_1[\mathcal{O}_2\mathcal{O}_2]_{n,l}} \sim \frac{1}{C_T} \sim \frac{1}{N^2}$$

(Hijano, Kraus, Perlmutter, Snively)

1-loop diagrams are further suppressed by $1/N^2 \sim 1/C_T \sim G_N/R^{d-1}$ where R is the AdS radius.

Introduction

In the high energy (eikonal) scattering graviton loops contribute powers of the ratio of Schwarzschild radius to the impact parameter, $(R_s/b)^k$, to the eikonal phase δ . $R_s/b \sim 1$ - black hole creation. (Amati, Ciafaloni, Veneziano)



Introduction

Making one particle heavy (black hole) allows one to compute δ classically, via computing the scattering angle.

At R_s/b increases to the critical value, δ becomes complex (absorption by the black hole).

Introduction

AdS-Schwarzschild background:

$$ds^2 = -f dt^2 + f^{-1} dr^2 + r^2 d\Omega_{d-1}^2,$$

where

$$f = 1 + \frac{r^2}{R^2} - \frac{\mu R^{d-2}}{r^{d-2}}$$

and

$$\mu \simeq \frac{G_N M}{R^{d-2}} \simeq \frac{\Delta_H}{C_T}$$

(by AdS/CFT dictionary $\Delta_H \simeq MR$ and $R^{d-1}/G_N \simeq C_T$). So $\Delta_H \sim C_T$, $\mu \simeq \Delta_H/C_T$ fixed is mapped to a dual black hole.

Introduction

In the CFT language, the stress-tensor contribution becomes

$$\lambda_{\mathcal{O}_H \mathcal{O}_H} T_{\mu\nu} \lambda_{\mathcal{O}_L \mathcal{O}_L} T_{\mu\nu} \sim \frac{\Delta_H}{\sqrt{C_T}} \frac{\Delta_L}{\sqrt{C_T}} \sim \mu$$

Double stress tensor contributions scale like

$$\lambda_{\mathcal{O}_H \mathcal{O}_H} [TT]_{n,\ell} \lambda_{\mathcal{O}_L \mathcal{O}_L} [TT]_{n,\ell} \sim \frac{\Delta_H^2}{C_T} \frac{\Delta_L^2}{C_T} \sim \mu^2$$

and k-stress tensor $[T_{\mu\nu}]^k$ contributes μ^k .

Introduction

Will be interested in the contribution of all operators $[T_{\mu\nu}]^k$ – the *stress tensor sector* of correlators. [Note: in $d = 2$ it goes under the name "the Virasoro vacuum block" and has been explicitly computed (Fitzpatrick, Kaplan, Walters).]

What about other operators (double trace, other light operators) ?

Few light operators in holographic theories (unless protected). Can decouple double trace operators by taking $\Delta_L \gg 1$. Also, some observables (like δ) are insensitive to double traces.

Outline

Leading twist multi stress tensors

- Double scaling lightcone limit

- Bootstrapping the near-lightcone correlator

Applications

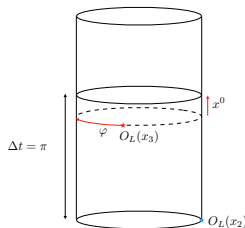
- Generalized Catalan numbers

- Holographic leading twist calculation

- Beyond leading twist

Double scaling lightcone limit

Correlator $\langle \mathcal{O}_H(t = +\infty) \mathcal{O}_L(\Delta t, \Delta\varphi) \mathcal{O}_L(0, 0) \mathcal{O}_H(t = -\infty) \rangle$ can be viewed as 2-point function $\langle \mathcal{O}_L \mathcal{O}_L \rangle_{\mathcal{O}_H}$ in the state created by \mathcal{O}_H at $t = \pm\infty$.



Cross ratios $z = e^{i(\frac{\Delta t}{R} + \Delta\varphi)}$, $\bar{z} = e^{i(\frac{\Delta t}{R} - \Delta\varphi)}$.

Double scaling lightcone limit

The stress-tensor exchange $\mathcal{O}_H \mathcal{O}_H - T_{\mu\nu} - \mathcal{O}_L \mathcal{O}_L$ produces $\mathcal{O}(\mu)$ term fixed by Ward identity; $\tau_1 = \ell(T_{\mu\nu}) - \Delta(T_{\mu\nu}) = d - 2$ (minimal twist for conserved current).

Correlator at $\mathcal{O}(\mu^2)$ is a result of an infinite sum over *double stress tensor operators*: $T_{\mu\nu} \partial_\alpha \dots \partial_\beta \square^n T_{\gamma\delta}$.

$$T_{\mu\nu} \partial_{\mu_1} \dots \partial_{\mu_s} T_{\alpha\beta}, \quad \tau_2 = 2(d - 2); \text{ leading twist}$$

$$T_\mu^\alpha T_{\alpha\nu}, \quad T_{\mu\nu} \partial_{\mu_1} \dots \partial_{\mu_s} \square T_{\alpha\beta}, \quad \tau = 2(d - 2) + 2$$

$$T_{\mu\nu} T^{\mu\nu}, \quad T_\mu^\alpha \square T_{\alpha\nu}, \quad T_{\mu\nu} \partial_{\mu_1} \dots \partial_{\mu_s} \square^2 T_{\alpha\beta}, \quad \tau = 2(d - 2) + 4$$

$\mathcal{O}(\mu^k)$ comes from k-stress tensors with leading twist $\tau_k = k(d - 2)$.

Double scaling lightcone limit

To compute the stress tensor sector, need to know all OPE coefficients. Things simplify in the lightcone limit $\bar{z} \rightarrow 1$. Lightcone OPE is dominated by lowest twist:

$$\mathcal{O}_1(x)\mathcal{O}_1(0) \sim x^{-2\Delta_1} \sum \frac{x^{\mu_1} \dots x^{\mu_\ell}}{(x^2)^{\frac{\tau}{2}}} \mathcal{O}_{\mu_1 \dots \mu_\ell}$$

This is reflected in the near-lightcone behavior of $\mathcal{B}_{[T_{\mu\nu}]^k} \sim (1 - \bar{z})^k \frac{d-2}{2}$ for leading twist $[T_{\mu\nu}]^k$. Their OPE coefficients are not affected by higher order gravitational terms (Fitzpatrick, Huang). Can keep all leading twist $[T_{\mu\nu}]^k$ by keeping $\tilde{\mu} = \mu (1 - \bar{z})^{\frac{d-2}{2}}$ fixed.

Bootstrapping the near-lightcone correlator

Ansatz + bootstrap in $d = 4$ (any even d works similarly)

$$\mathcal{O}(\tilde{\mu}) : \langle \mathcal{O}_H \mathcal{O}_H \mathcal{O}_L \mathcal{O}_L \rangle^{(1)} \simeq [(1-z)(1-\bar{z})]^{-\Delta_L} f_3(z)$$

where

$$f_a(z) = (1-z)^a {}_2F_1(a, a, 2a, 1-z)$$

At higher orders, make an educated guess (ansatz):

$$\mathcal{O}(\tilde{\mu}^2) : \langle \mathcal{O}_H \mathcal{O}_H \mathcal{O}_L \mathcal{O}_L \rangle^{(2)} \sim b_{33} f_3^2 + b_{24} f_2 f_4 + b_{15} f_1 f_5$$

$$\mathcal{O}(\tilde{\mu}^3) : \langle \mathcal{O}_H \mathcal{O}_H \mathcal{O}_L \mathcal{O}_L \rangle^{(3)} \sim b_{333} f_3^3 + b_{234} f_2 f_3 f_4 + \dots$$

and similarly for all values of k for $\mathcal{O}(\tilde{\mu}^k)$ coefficients. Can bootstrap and solve for all b_i 's, order by order.

Bootstrapping the near-lightcone correlator

Consider stress-tensor contribution, $k = 1$ term in the $\mathcal{O}_H \mathcal{O}_H - \mathcal{O}_L \mathcal{O}_L$ channel:

$$\langle \mathcal{O}_H \mathcal{O}_H \mathcal{O}_L \mathcal{O}_L \rangle^{(1)} \sim (\alpha_0 + \alpha_1 z + \dots) + (\beta_0 + \beta_1 z + \dots) \log z$$

with α_i, β_i known coefficients. The cross channel, $\mathcal{O}_H \mathcal{O}_L - \mathcal{O}_H \mathcal{O}_L$, produces

$$\langle \mathcal{O}_H \mathcal{O}_H \mathcal{O}_L \mathcal{O}_L \rangle^{(1)} \sim (P_0^{(1)} + P_1^{(1)} z + \dots) + (\gamma_0^{(1)} + \gamma_1^{(1)} z + \dots) \log z$$

where $\gamma(n, \ell) = \gamma_n^{(1)}/\ell + \mu^2 \gamma_n^{(2)}/\ell^2 + \dots$ and similarly $P(n, \ell)$ are anomalous dimensions and OPE coefficients of heavy-light double trace operators $[\mathcal{O}_H \mathcal{O}_L]_{n, \ell}$.

Bootstrapping the near-lightcone correlator

This determines $P_n^{(1)}, \gamma_n^{(1)}$ and all terms $\sim \log^2 z$ at $\mathcal{O}(\tilde{\mu}^2)$. This determines **all** coefficients in the ansatz at $\mathcal{O}(\tilde{\mu}^2)$:

$$\begin{aligned} \langle \mathcal{O}_H \mathcal{O}_H \mathcal{O}_L \mathcal{O}_L \rangle^{(2)} \simeq & \frac{1}{[(1-z)(1-\bar{z})]^{\Delta_L}} \left(\frac{\Delta_L}{\Delta_L - 2} \right) \times \\ & \left[(\Delta_L - 4)(\Delta_L - 3)f_3(z)^2 + \frac{15}{7}(\Delta_L - 8)f_2(z)f_4(z) \right. \\ & \left. + \frac{40}{7}(\Delta_L + 1)f_1(z)f_5(z) \right] \end{aligned}$$

Bootstrapping the near-lightcone correlator

OPE coefficients with the leading twist double-stress operators:

$$\lambda_{\mathcal{O}_H \mathcal{O}_H [T_{\mu\nu}^2]_{n=0,\ell}} \lambda_{\mathcal{O}_L \mathcal{O}_L [T_{\mu\nu}^2]_{n=0,\ell}} = \frac{\Delta_H^2}{C_T} \frac{a_\ell^2}{C_T} \frac{\Delta_L}{\Delta_L - 2} (\Delta_L^2 + b_\ell \Delta_L + c_\ell)$$

where

$$b_\ell = -1 + \frac{36}{\ell(\ell+3) + c_\ell}, \quad c_\ell = \frac{288}{(\ell-2)\ell(\ell+3)(\ell+5)}, \quad a_\ell = \dots$$

so

$$\lambda_{\mathcal{O}_L \mathcal{O}_L [T_{\mu\nu}^2]_{n=0,\ell}} = \frac{a_\ell}{C_T} \frac{\Delta_L}{\Delta_L - 2} (\Delta_L^2 + b_\ell \Delta_L + c_\ell)$$

Bootstrapping the near-lightcone correlator

One can continue this procedure and compute the lightcone correlator to any desired order in $\tilde{\mu}$

We can read off the OPE coefficients $\lambda_{\mathcal{O}_L \mathcal{O}_L}[T_{\mu\nu}^k]_{n=0,\ell}$ to leading order in $1/C_T$. We did not use holography, but recover the OPE coefficients which have been computed using holography (and get many more).

The correlator exponentiates

$$\langle \mathcal{O}_H \mathcal{O}_H \mathcal{O}_L \mathcal{O}_L \rangle \sim e^{\Delta_L \mathcal{F}(z, \tilde{\mu}, \Delta_L)}$$

where $\mathcal{F}(z, \tilde{\mu}, \Delta_L)$ has a finite limit \mathcal{F}_∞ as $\Delta_L \rightarrow \infty$.

Bootstrapping the near-lightcone correlator

The result further simplifies in the large volume limit ($R \rightarrow \infty$). Only operators $T_{\mu\nu} \dots T_{\alpha\beta}$ contribute (no derivatives allowed) and the result is an expansion in powers of $\mu \Delta x^- (\Delta x^+)^3$:

$$\begin{aligned} \mathcal{F}_\infty|_{d=4} = & -\log(\Delta x^+ \Delta x^-) + \frac{\mu \Delta x^- (\Delta x^+)^3}{120} \\ & + \frac{\mu^2 (\Delta x^-)^2 (\Delta x^+)^6}{10080} + \frac{1583 \mu^3 (\Delta x^-)^3 (\Delta x^+)^9}{648648000} + \dots, \end{aligned}$$

Generalized Catalan numbers

In $d = 2$ the HHLL Virasoro vacuum block satisfies

$$\mathcal{F}_{Vir} \simeq B_2(\mu) \log z, \quad z \rightarrow 0$$

where $B_2(\mu) = \frac{1}{2}(1 - \sqrt{1 - 4\mu})$, the solution of

$$B_2(\mu) = B_2^2(\mu) + \mu$$

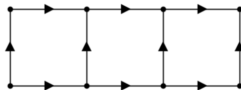
is the generating function of the Catalan numbers

$C_n = 1, 2, 5, 14, 42, \dots$,

$$B_2(\mu) = \sum_{n=1}^{\infty} C_n \mu^n, \quad C_n = \frac{(2n)!}{(n+1)!n!}$$

Catalan numbers

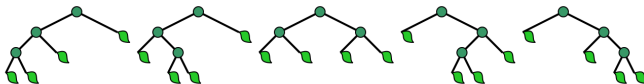
The number of linear extensions of a partially ordered set (poset):



Number of $2 \times n$ Young tableaux

4 5 6	3 5 6	3 4 6	2 5 6	2 4 6
1 2 3	1 2 4	1 2 5	1 3 4	1 3 5

Number of binary trees with $n + 1$ leaves



Generalized Catalan numbers

The equation for the generating function $B_2(\mu)$ can be generalized to a nonlinear differential equation that \mathcal{F}_{Vir} satisfies,

$$\frac{d^2 \mathcal{F}_{Vir}(\mu, z)}{dz^2} = \left(\frac{d \mathcal{F}_{Vir}(\mu, z)}{dz} \right)^2 + \frac{\mu}{z^2}$$

(Fitzpatrick, Kaplan, Walters, Wang). The solution satisfies the ansatz

$$\mathcal{F}_{Vir}(\mu, z) = a_2 \mu f_2 + \mu^2 (a_{22} f_2 f_2 + a_{13} f_1 f_3) + \dots$$

Generalized Catalan numbers

What about higher spin algebras? Consider HHLL blocks of W_3 with $\mu \simeq Q_3/c$ finite. It satisfies the ansatz

$$\mathcal{F}_{W_3}(\mu, z) = a_3 \mu f_3 + \mu^2 (a_{33} f_3 f_3 + a_{24} f_2 f_4 + a_{15} f_1 f_5) + \dots$$

and, moreover,

$$\mathcal{F}_{W_3} \simeq B_3(\mu) \log z + 0 \cdot \log^n z, \quad z \rightarrow 0$$

For $h = 3q_3$, $B_3(\mu)$ satisfies a generalization of the Catalan equation

$$B_3(\mu) = -2B_3(\mu)^2 + 3B_3(\mu)^3 + \mu$$

Generalized Catalan numbers

In fact, $\mathcal{F}_{W3}(\mu, z)$ satisfies a differential equation

$$\frac{d^3 \mathcal{F}_{W3}}{dz^3} = -2 \left(\frac{d^2 \mathcal{F}_{W3}}{dz^2} \right) \left(\frac{d \mathcal{F}_{W3}}{dz} \right) + 3 \left(\frac{d \mathcal{F}_{W3}}{dz} \right)^3 + \frac{\mu}{z^2}$$

Similar patterns are observed for higher W_N algebras.

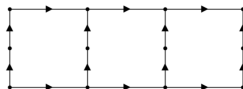
Generalized Catalan numbers

What about the near lightcone $d = 4$ CFT result? It satisfies

$$\mathcal{F} \simeq B(\mu) \log z + 0 \cdot \log^n z, \quad z \rightarrow 0$$

in the $\Delta_L \rightarrow 0$ limit.

The numbers that appear in $B(\mu) = \sum \tilde{C}_n \mu^n$ are $\tilde{C}_n = 1, 6, 71, 1266, \dots$. These are the numbers of linear extensions of the poset



We don't yet know an equation for their generating function.

Holographic leading twist calculation

In holographic CFTs we can compute

$$\langle \mathcal{O}_H(t = +\infty) \mathcal{O}_L(\Delta t, \Delta\varphi) \mathcal{O}_L(0, 0) \mathcal{O}_H(t = -\infty) \rangle$$

for $\Delta_L \gg 1$ by computing

$$\langle \mathcal{O}_L(\Delta t, \Delta\varphi) \mathcal{O}_L(0, 0) \rangle = e^{-\Delta_L \ell}$$

where $\ell = -\mathcal{F}_\infty$ is the regularized length of the geodesic which connects points $(\Delta t, \Delta\varphi)$ and $(0, 0)$ on the boundary of AdS-Schwarzschild spacetime.

Holographic leading twist calculation

Rescale coordinates in the AdS-Schwarzschild metric,
 $x^- = (t - \varphi)\mu^{\frac{2}{d-2}}$ and $y = r\mu^{-\frac{1}{d-2}}$ and consider the $\mu \rightarrow \infty$ limit
 keeping x^\pm , y fixed (here $R = 1$)

$$ds^2 = -\frac{1}{4} \left(1 - \frac{1}{y^{d-2}}\right) (dx^+)^2 - y^2 dx^+ dx^- + \frac{dy^2}{y^2}.$$

Two Killing vectors, ∂_+ and ∂_- give rise to two conserved quantities, K and K_+ . Geodesic equation (spacelike) becomes

$$\dot{y}^2 + 4KK_+ + (y^{-2} - y^{-d})K^2 - y^2 = 0.$$

Holographic leading twist calculation

Can find the length of the geodesic $\ell(\Delta x^+, \Delta x^-)$. $\mathcal{F}_\infty = -\ell$ but it's UV-divergent, need to regularize. Things simplify in the large volume limit.

The solution can be written in terms of elliptic integrals [Appel functions]

$$\ell_f \simeq \log(\Delta x^+ \Delta x^-) - \log[I_+(\alpha)I_-(\alpha)] + I_\ell(\alpha)$$

where α is determined by

$$(-\Delta x^-)^{\frac{d-2}{2}} (\Delta x^+)^{\frac{d+2}{2}} = \alpha I_-^{\frac{d-2}{2}}(\alpha) I_+^{\frac{d+2}{2}}(\alpha)$$

and

$$I_\ell(x) = 2 \int_{u_0}^{\Lambda_u} \frac{u^{\frac{d}{2}} du}{(u^{d+2} - 4u^d + x)^{\frac{1}{2}}} - 2 \log \Lambda_u$$

Holographic leading twist calculation

$d = 2$: recover HHLL Virasoro block:

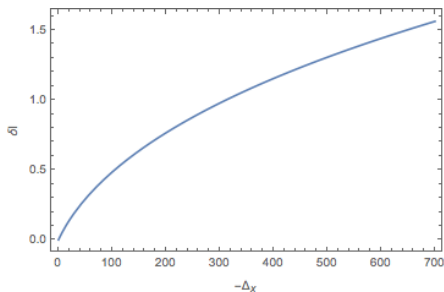
$$-\ell_f|_{d=2} \simeq -\log \sinh \frac{\sqrt{\mu} \Delta_x^+}{2},$$

$d = 4$: new result; expansion in

$\Delta_x = \mu \Delta_x^- (\Delta_x^+)^3 \simeq T^4 \Delta_x^- (\Delta_x^+)^3$ (where T is the temperature). Agrees with bootstrap:

$$\begin{aligned} -\ell_f|_{d=4} \simeq & -\log(\Delta_x^- \Delta_x^+) + \frac{\Delta_x}{120} + \frac{\Delta_x^2}{10080} + \frac{1583 \Delta_x^3}{648648000} \\ & + \frac{3975313 \Delta_x^4}{49401031680000} + \dots \end{aligned}$$

Holographic leading twist calculation



$\mathcal{F}_\infty|_{d=4}$ with $\log(\Delta x^+ \Delta x^-)$ term subtracted as a function of $\mu \Delta x^- (\Delta x^+)^3$.

Beyond leading twist

Bootstrap can be successfully extended beyond the leading twist.

The results agree with the phase shift (a.k.a. eikonal phase) δ , which has been explicitly computed in gravity to all orders in the impact parameter for the HHLL case.

What happens in non-holographic theories? There are finite gap corrections to $\lambda_{\mathcal{O}\mathcal{O}[T_{\mu\nu}]^k}$ (Fitzpatrick, Huang, Meltzer, Perlmutter, Simmons-Duffin). Is the stress tensor sector still special in any way?

Thermalization of stress-tensor sector

In large-N CFTs stress-tensor thermalizes,

$$\langle \mathcal{O}_H | T_{\mu\nu} | \mathcal{O}_H \rangle = \langle T_{\mu\nu} \rangle_\beta$$

where β is determined by the saddle-point equation,

$\Delta_H/R = \partial_\beta(\beta F)$. Large-N factorization implies

$\langle [T_{\mu\nu}]^k \rangle_\beta = \langle T_{\mu\nu} \rangle_\beta^k \simeq \Delta_H^k$. Hence, thermalization of $[T_{\mu\nu}]^k$ is equivalent to

$$\langle \mathcal{O}_H | [T_{\mu\nu}]^k | \mathcal{O}_H \rangle = \# \Delta_H^k \left[1 + \mathcal{O}\left(\frac{1}{C_T}\right) \right]$$

This is a statement about the leading Δ behavior of multi-stress tensor OPE coefficients, which can be verified.

Summary

- ▶ Can compute HHLL correlator to all orders in μ via bootstrap.
- ▶ Form of \mathcal{F} in CFT_d is the same as that for $W_{d/2+1}$ algebras in $d = 2$. Can write equation for the generating function in the W_N algebra case.
- ▶ Computed it in the large Δ_L limit in holography – perfect agreement.
- ▶ Stress tensor sector thermalizes (even in non-holographic large N CFTs).

To do list

- ▶ Find closed form of heavy-heavy-light-light correlator.
Generating function?
- ▶ Symmetry of the lightcone correlator?
- ▶ Understand thermal strongly coupled large-N CFTs
- ▶ Regge scattering in gravity

Thank you!